



DM Predictions from the LHC

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Corpus Cristi, 05/2019

1. Introduction & Models
2. Results in SUSY GUT models
3. Results in the pMSSM11
4. Results in non-SUSY models
5. Conclusions



1. Introduction & Models

GUT based models:

- 1.) CMSSM: $m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$
- 2.) NUHM1: CMSSM + 1 scalar mass parameter
 $m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$ and M_A
- 3.) NUHM2: CMSSM + 2 scalar mass parameters
 $m_0, m_{1/2}, A_0, \tan \beta, \mu$ and M_A
- 4.) SU(5): CMSSM + 3 scalar mass parameters
 $m_5, m_{10}, m_{1/2}, A_0, \tan \beta, m_{H_u}, m_{H_d}$
- 5.) mAMSB: different mechanism for SUSY breaking
 $m_{3/2}, m_0, \tan \beta, \text{sign}(\mu)$
- 6.) sub-GUT: CMSSM, but unification at lower scale
 $m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$ and M_{in}
- 7.) ...

⇒ wide variety of models covered!

Problem: We cannot be sure about the SUSY-breaking mechanism

- ⇒ it is possible that with the CMSSM, NUHM, SU(5), mAMSB, sub-GUT we missed the “correct” mechanism
- ⇒ hint: strong connection between colored and uncolored sector
tension between low-energy EW effects and (colored) LHC searches

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tension between low-energy EW effects and (colored) LHC searches

Solution: investigate also the “general MSSM”

⇒ 11 parameters are manageable ⇒ pMSSM11

- squark mass parameters: $m_{\tilde{q}_{1,2}} =: m_{\tilde{q}}, m_{\tilde{q}_3}$
- slepton mass parameter(s): $m_{\tilde{l}}, m_{\tilde{\tau}}$
- gaugino masses: M_1, M_2, M_3
- trilinear coupling: A
- Higgs sector parameters: $M_A, \tan \beta$
- Higgs mixing parameter: μ

What if we still did not get it right?

- low-energy model different?
- richer SUSY structure?
- no SUSY model? \Rightarrow not really realistic!

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Lagrangian according to LHC-DM-WG recommendation:

The Lagrangians

- We consider DMSMs with a spin-1 (Y_1) s-channel mediator.
- The dark matter candidate is a Dirac fermion (X_D).
- We use the model files provided by the DMSIMP package for our implementation.

Spin-1 mediator

- Interaction Lagrangian mediator-DM
$$\mathcal{L}_{X_D}^{Y_1} = \bar{X}_D \gamma_\mu \left(g_{X_D}^V + g_{X_D}^A \gamma_5 \right) X_D Y_1^\mu.$$
- Interaction Lagrangian mediator-quarks
$$\mathcal{L}_{quarks}^{Y_1} = \sum_{i,j} \left[\bar{d}_i \gamma_\mu \left(g_{d_{i,j}}^V + g_{d_{i,j}}^A \gamma_5 \right) d_j + \bar{u}_i \gamma_\mu \left(g_{u_{i,j}}^V + g_{u_{i,j}}^A \gamma_5 \right) u_j \right] Y_1^\mu$$
- Interaction Lagrangian mediator-leptons
$$\mathcal{L}_{leptons}^{Y_1} = \sum_{i,j} \left[\bar{l}_i \gamma_\mu \left(g_{l_{i,j}}^V + g_{l_{i,j}}^A \gamma_5 \right) l_j \right] Y_1^\mu$$

Scenarios

- Leptophobic, $g_{l_{i,j}}^V = g_{l_{i,j}}^A = 0$ (no constraints from dilepton searches).
- Flavor diagonal, $g_{u/d_{i,j}}^{V/A} = 0$ if $i \neq j$.
- Flavor blind, $g_{u_{i,j}}^{V/A} = g_{d_{i,j}}^{V/A}$.

1. $g_{X_D}^V \equiv g_{DM}$ $g_{X_D}^A = 0$
 $g_{u/d}^V \equiv g_{SM}$ $g_{u/d}^A = 0$,
pure vector.
2. $g_{X_D}^V = 0$ $g_{X_D}^A \equiv g_{DM}$
 $g_{u/d}^V = 0$ $g_{u/d}^A = g_{SM}$,
pure axial-vector.

[taken from E. Bagnaschi]

Our tool: **Mastercode**



⇒ collaborative effort of theorists and experimentalists

[Bagnaschi, Borsato, Buchmüller, Chobanova, Citron, Costa, De Roeck, Dolan, Ellis, Flächen, Hahn, SH, Isidori, Lucio, Martinez Santos, Olive, Trifa, Sakurai, Weiglein]

Über-code for the combination of different tools:

- Über-code original in Fortran, now re-written in C++
- tools are included as **subroutines**
- **compatibility** ensured by collaboration of authors of “MasterCode” and authors of “sub tools” **/SLHA(2)**
- sub-codes in Fortran or C++

⇒ evaluate observables of one parameter point consistently with various tools

cern.ch/mastercode

Data we have:

- Higgs boson mass/couplings/... (LHC) \Rightarrow FeynHiggs

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- SUSY / di-jet /mono-jet searches (LHC) \Rightarrow own re-cast

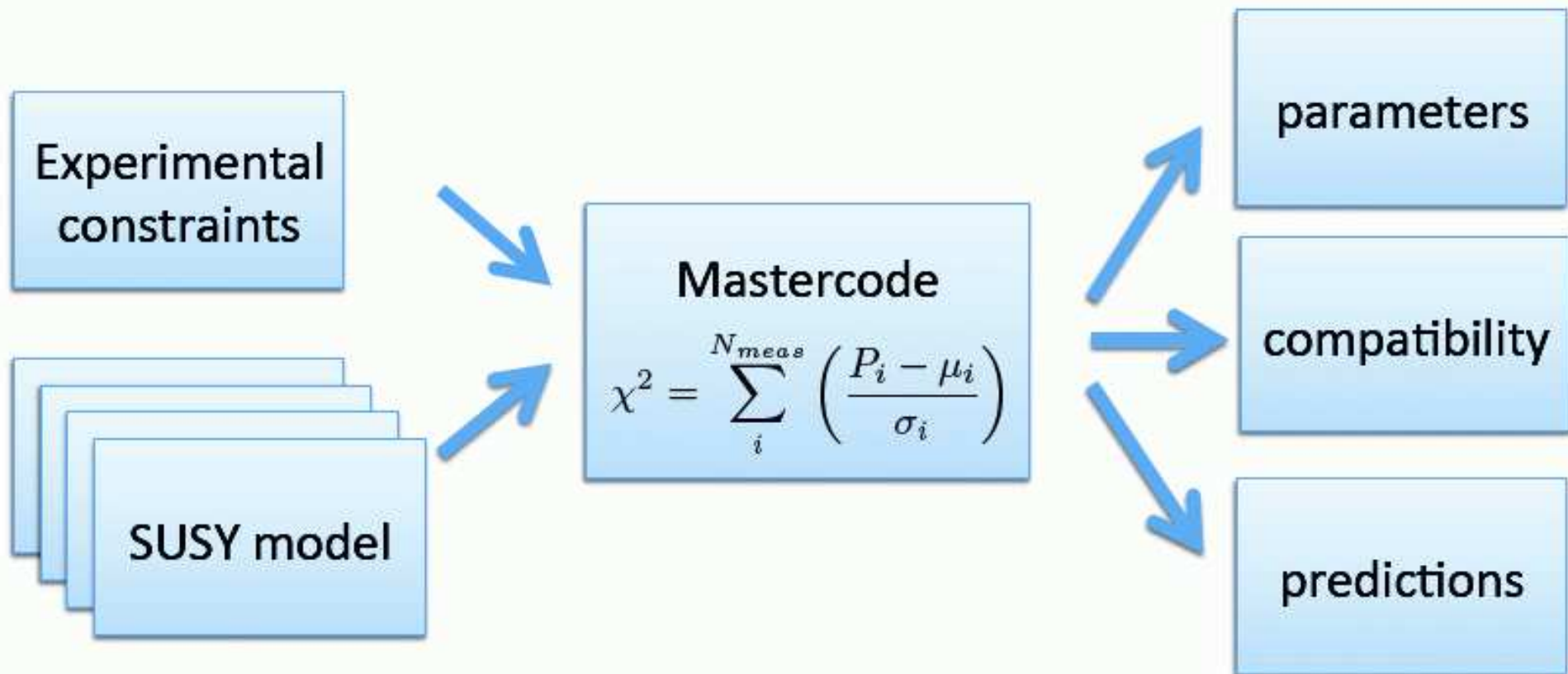
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- SUSY / di-jet /mono-jet searches (LHC) \Rightarrow own re-cast
- electroweak precision data \Rightarrow FeynWZ, FeynHiggs
- flavor data \Rightarrow SuperIso, SuFla
- astrophysical data (DM properties) \Rightarrow MicrOMEGAs, SSARD

The χ^2 evaluation:

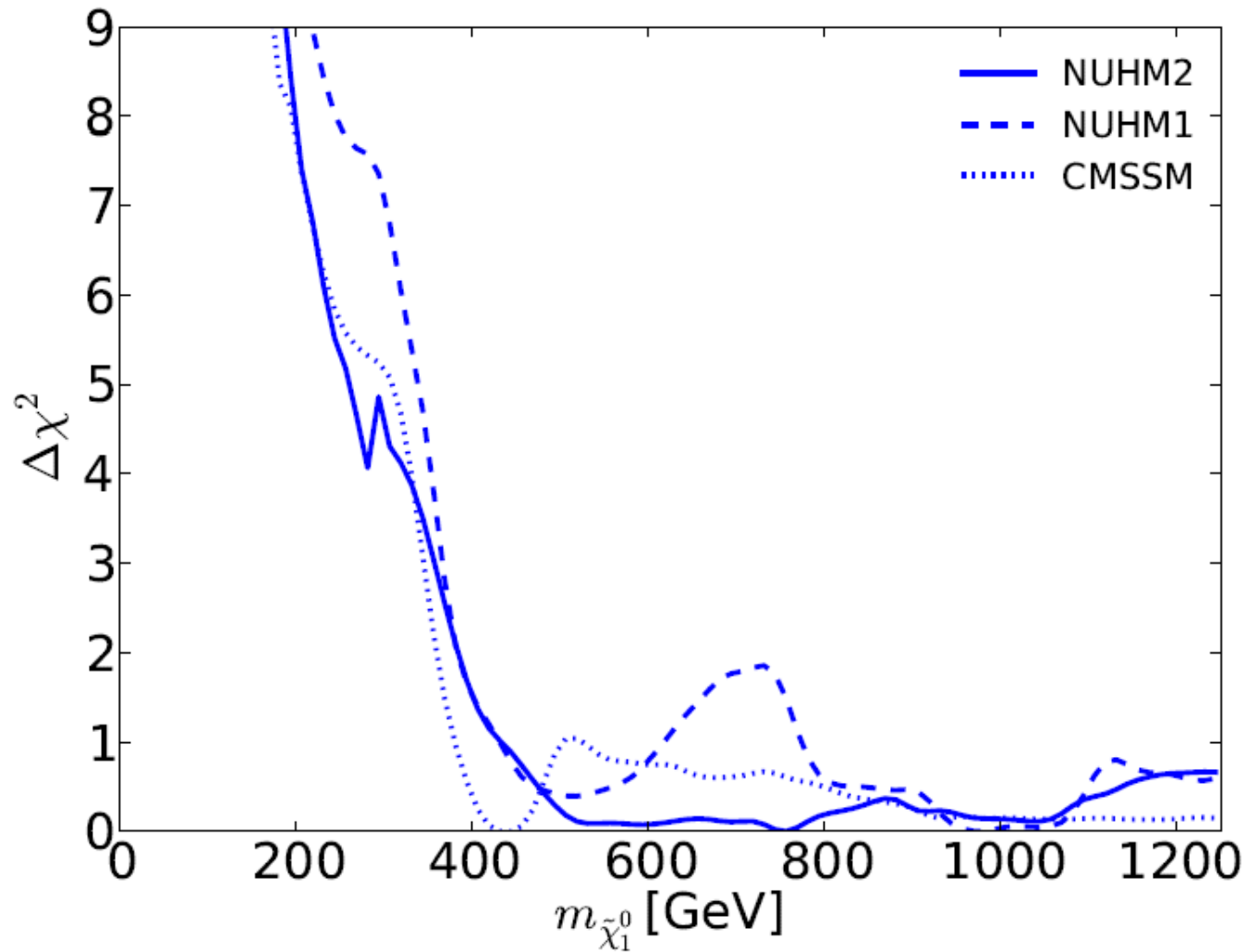


Global fits of SUSY

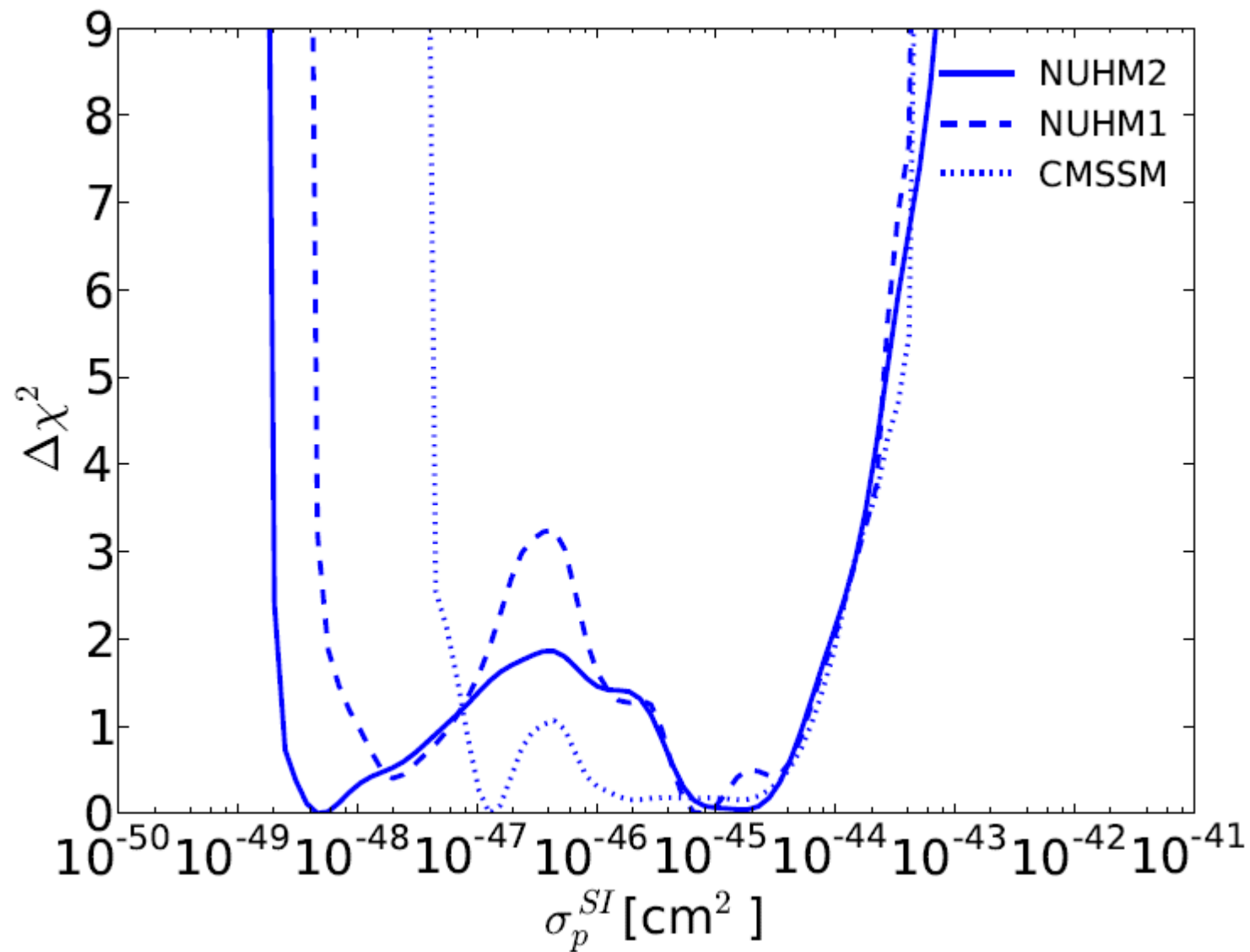


2. Results in SUSY GUT models

Results in the CMSSM, NUHM1, NUHM2

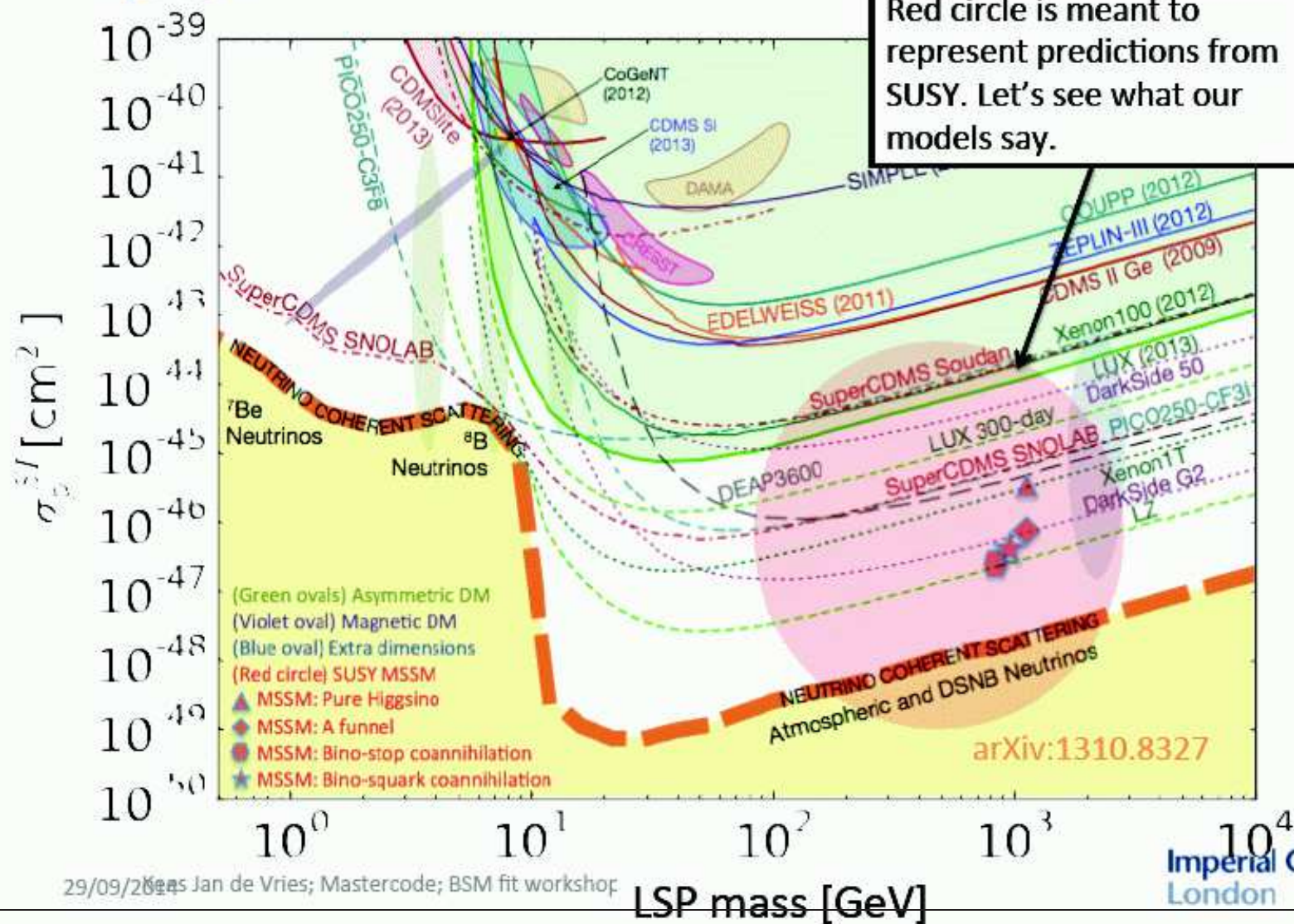


⇒ only very large values are favored



⇒ only very small values are favored

mastercode direct detection: past-present-future

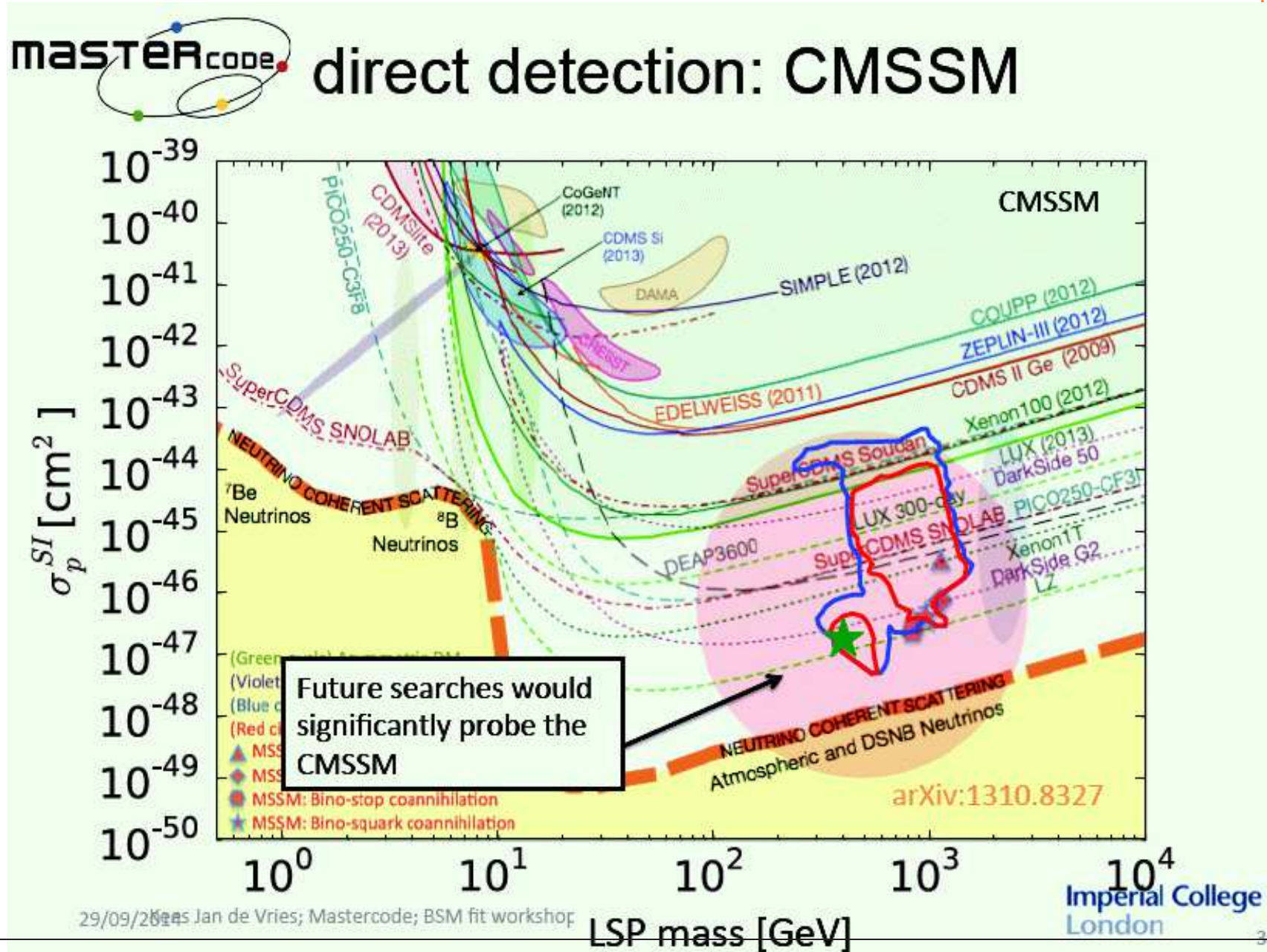


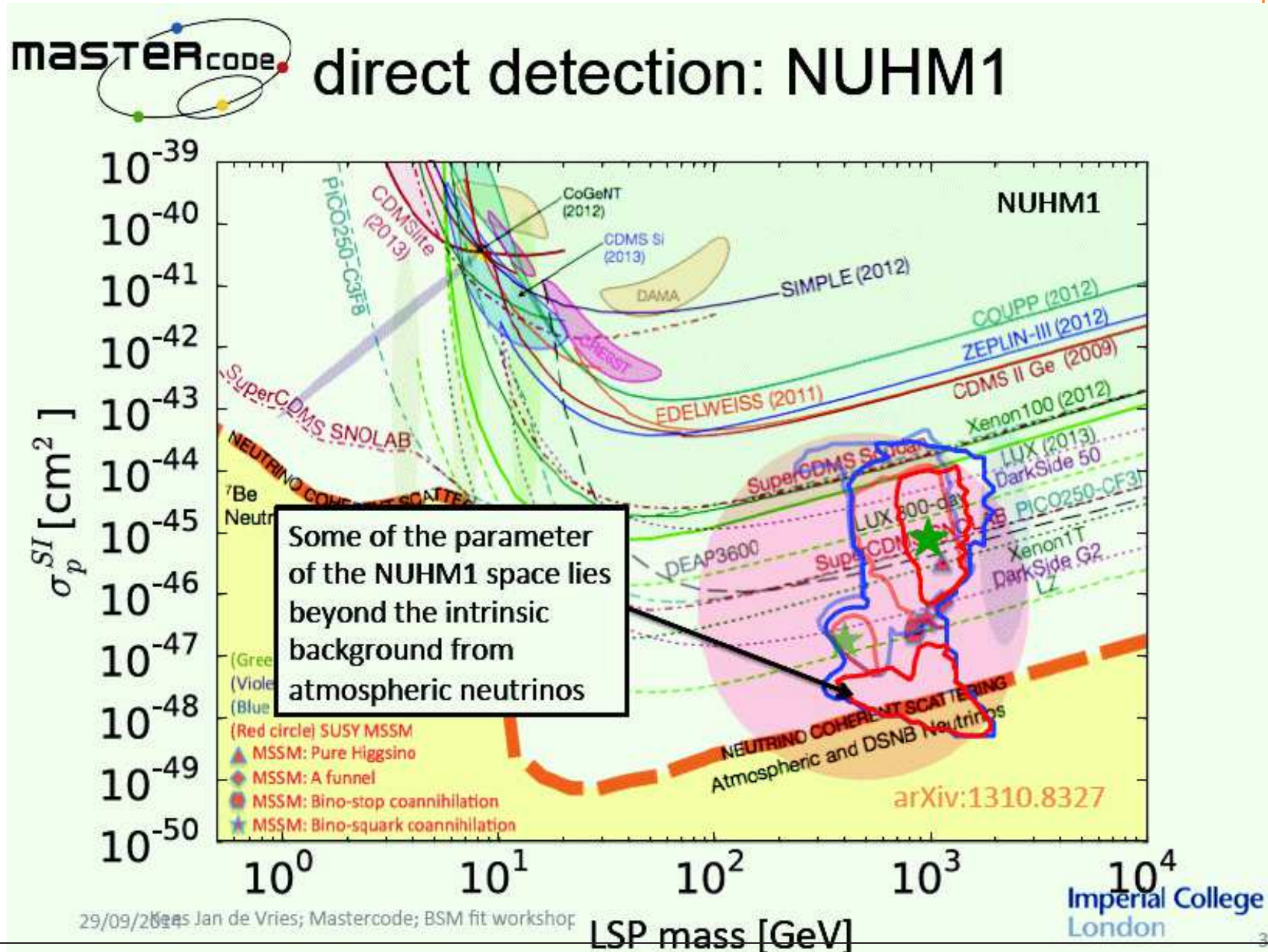
29/09/2014 Jan de Vries; Mastercode; BSM fit workshop

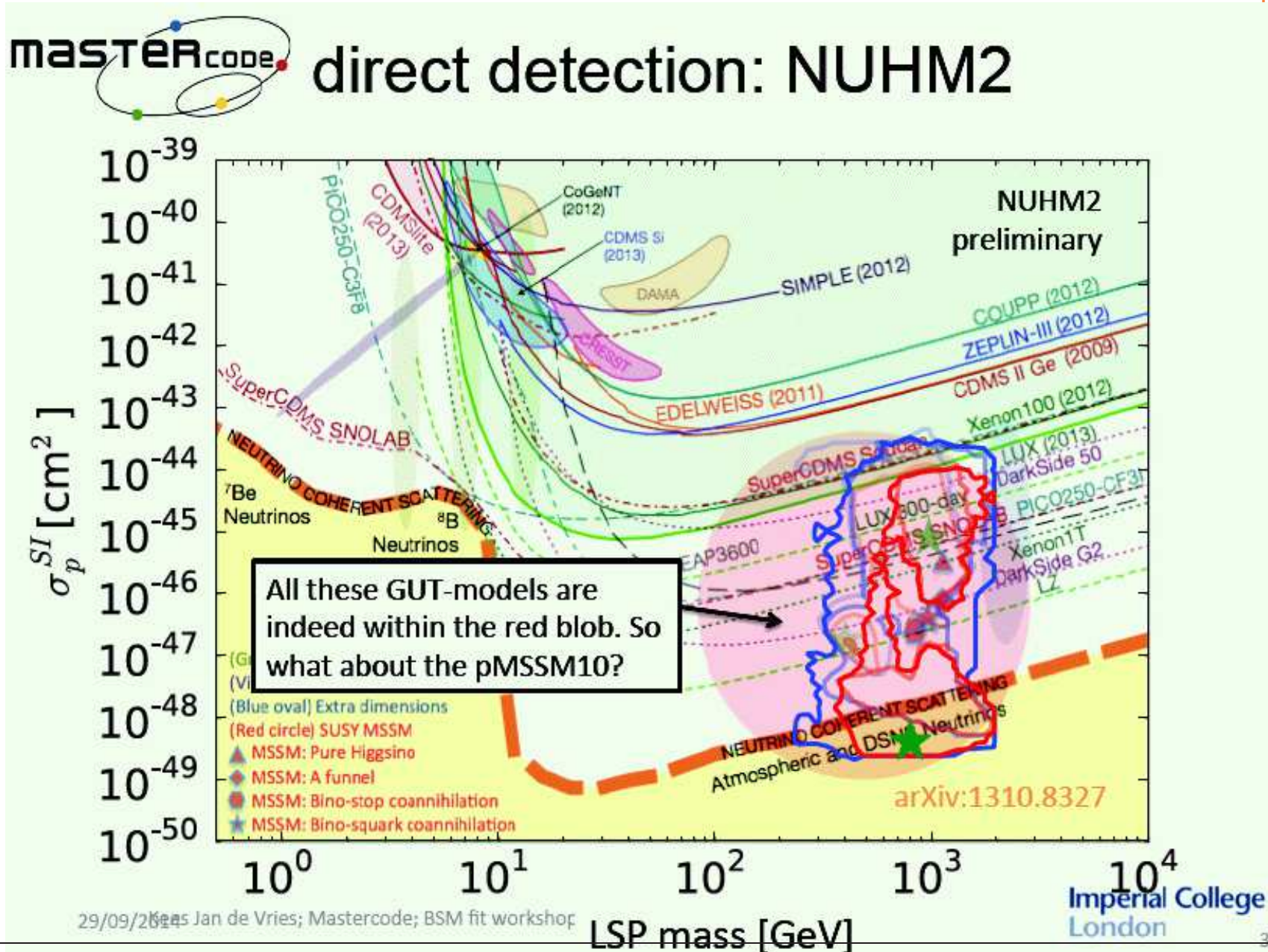
LSP mass [GeV]

Imperial College
London

29





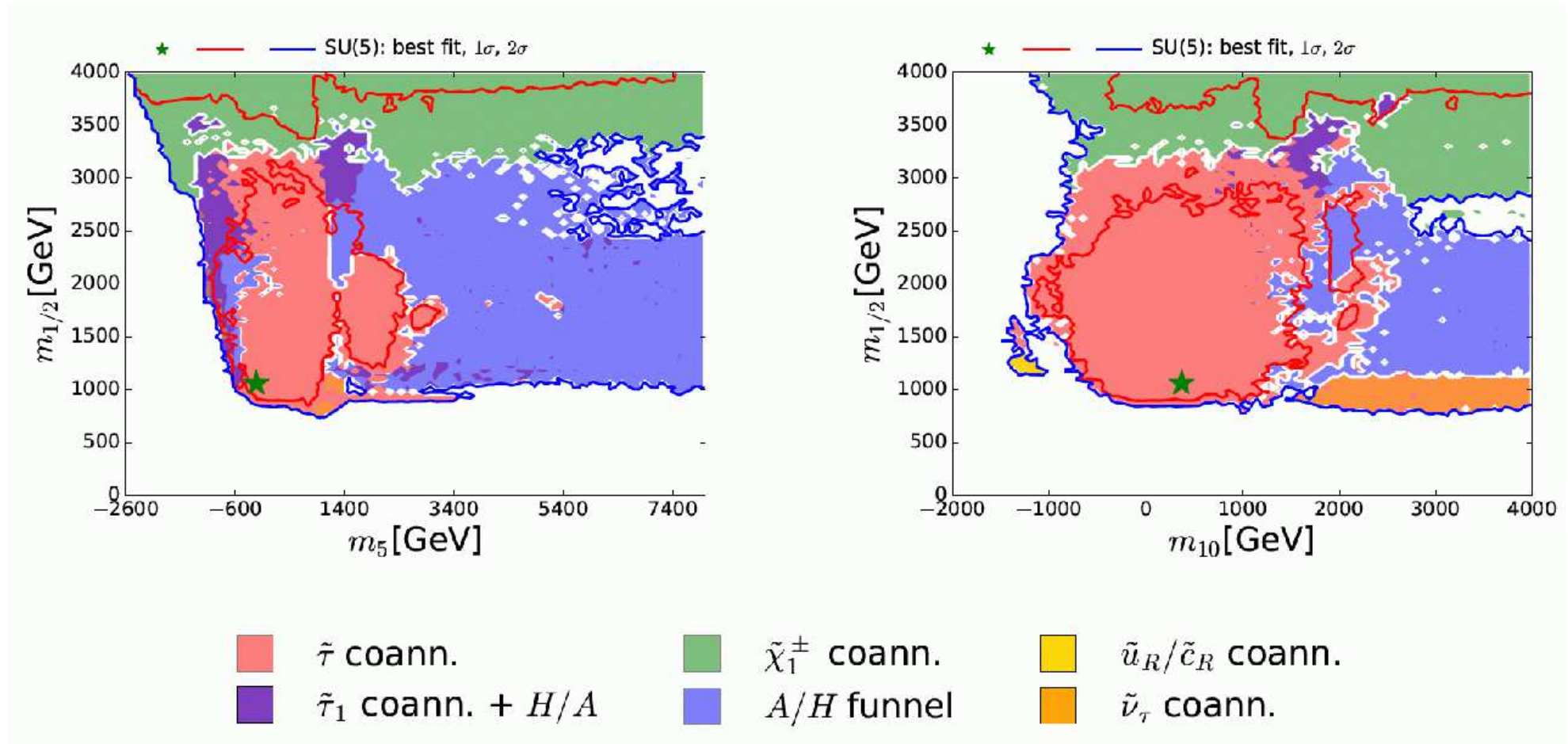


Results in the SU(5)



Dark Matter annihilation mechanism:

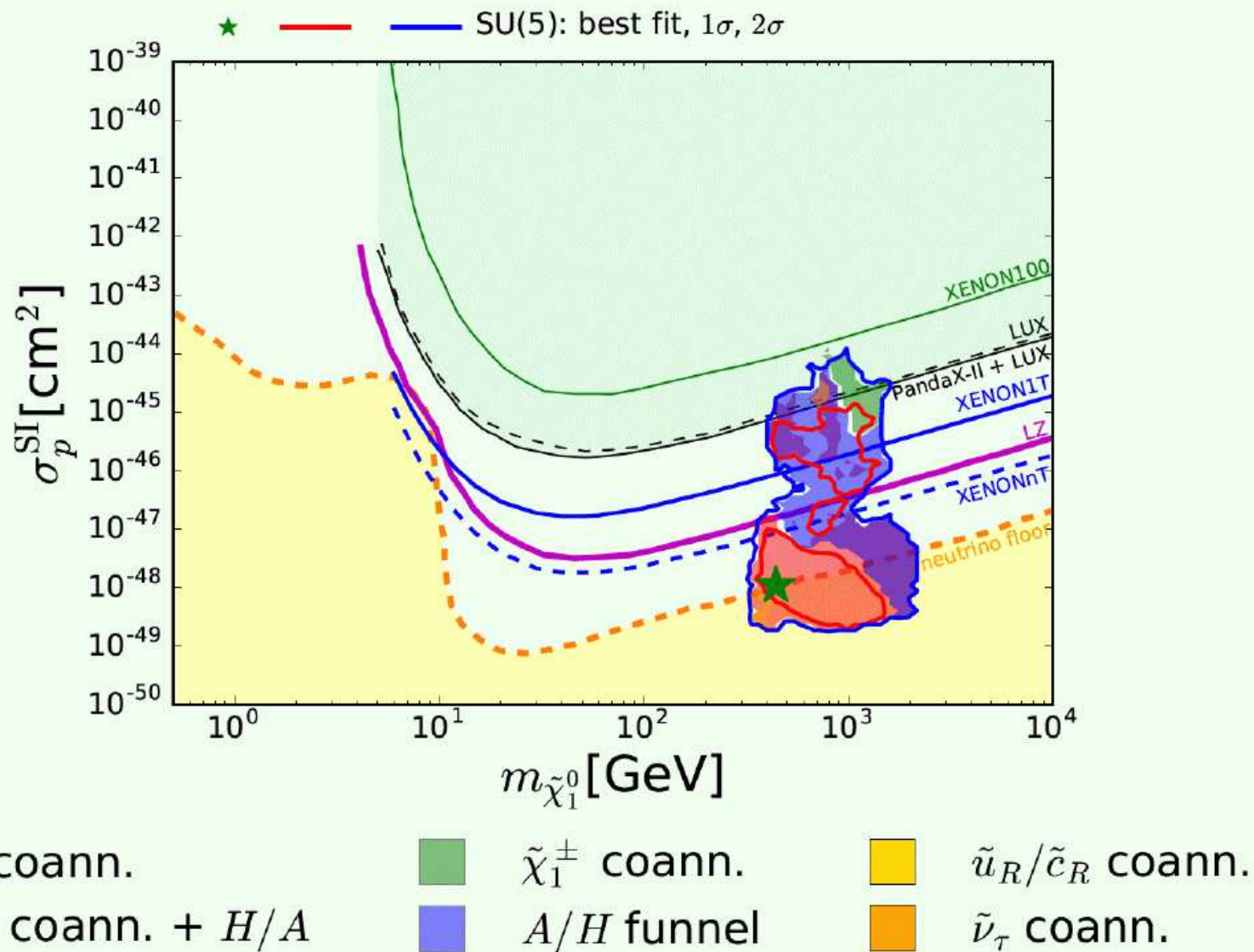
[2016]



$\Rightarrow \tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$ co-annihilation possible

Dark Matter Direct Detection prospects:

[2016]

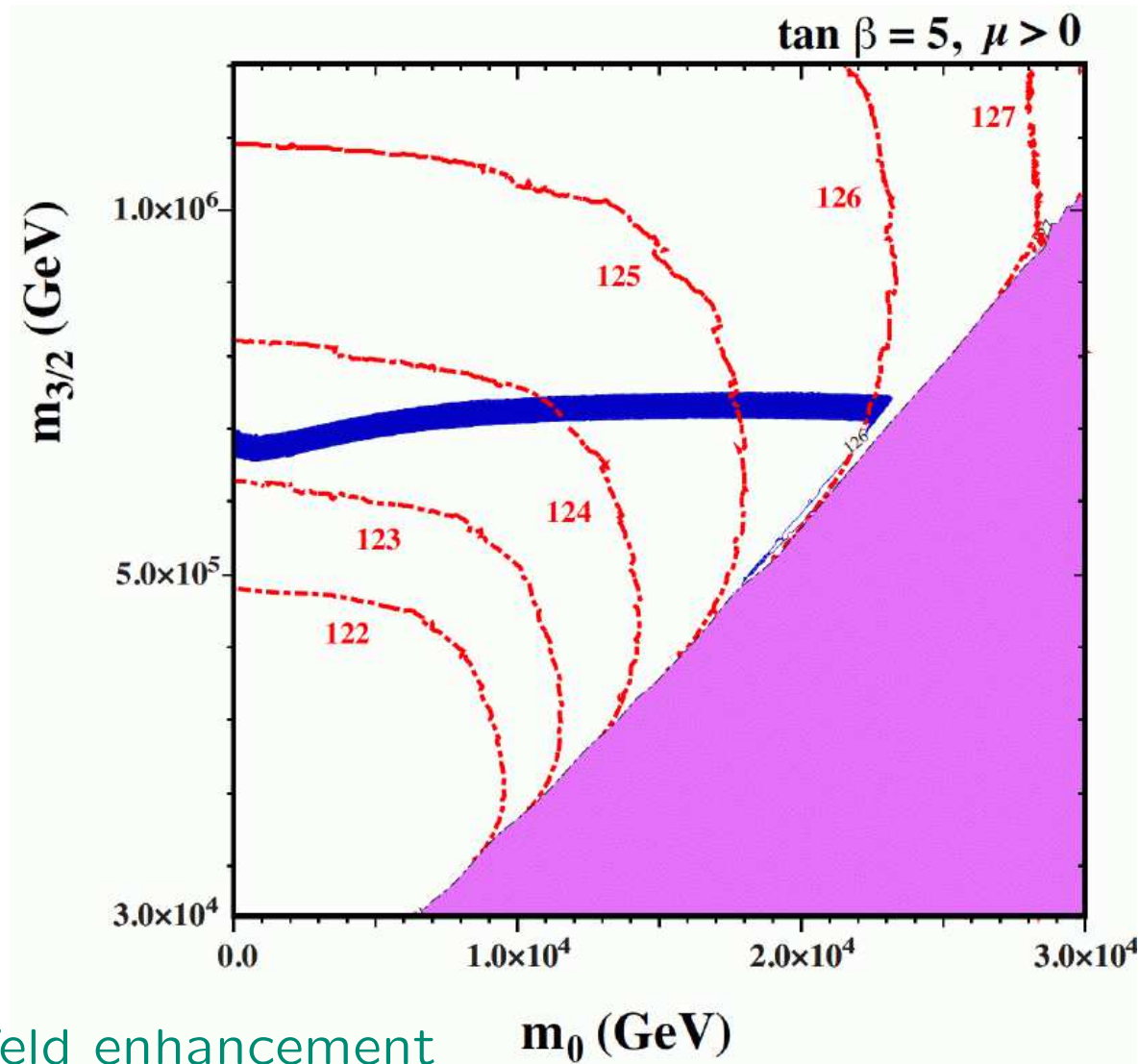


Results in the mAMSB



Known fact: Dark Matter requirement restricts $m_{3/2}$:

[2016]



⇒ no Sommerfeld enhancement

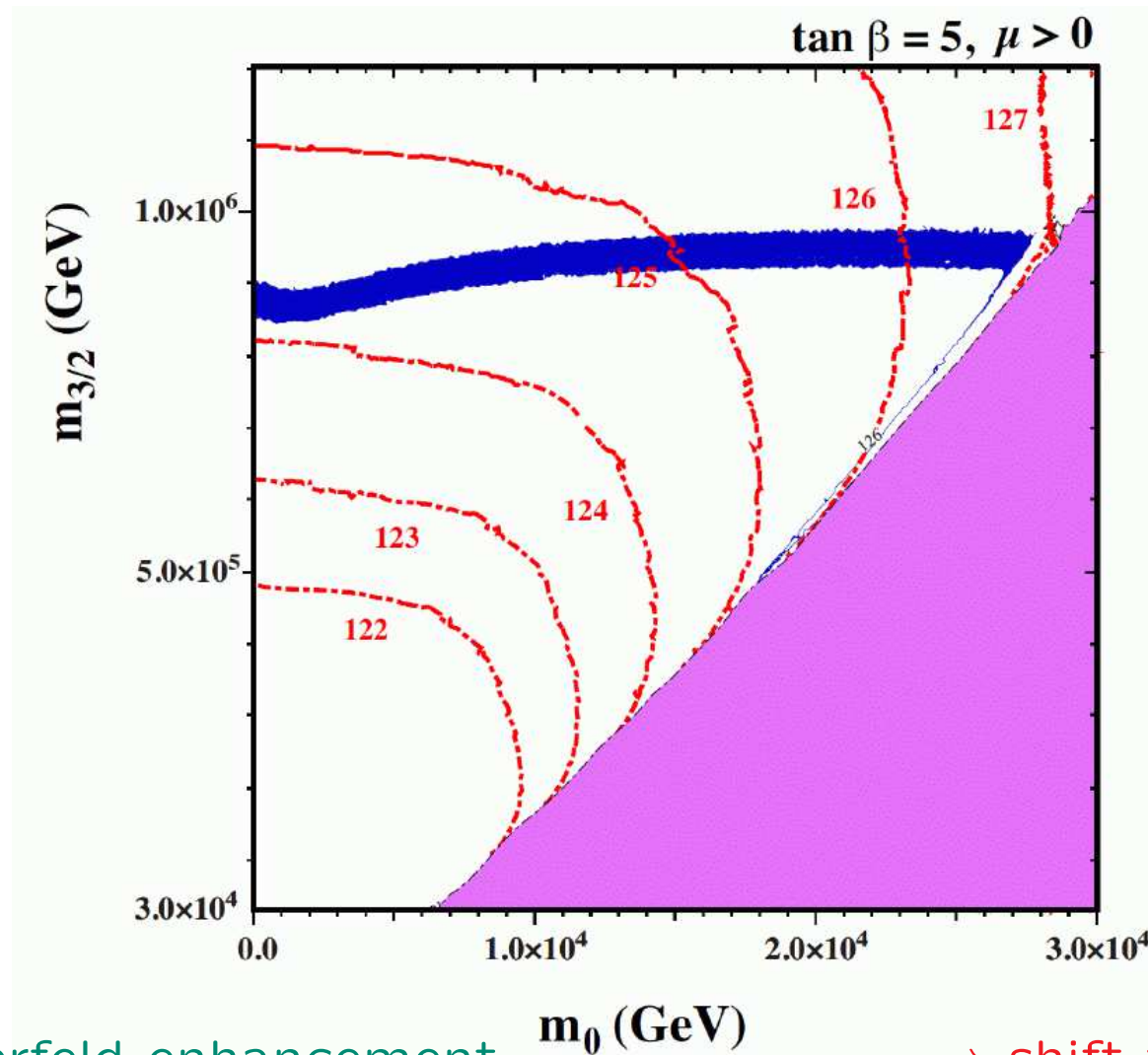
m_0 (GeV)

Results in the mAMSB



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[2016]

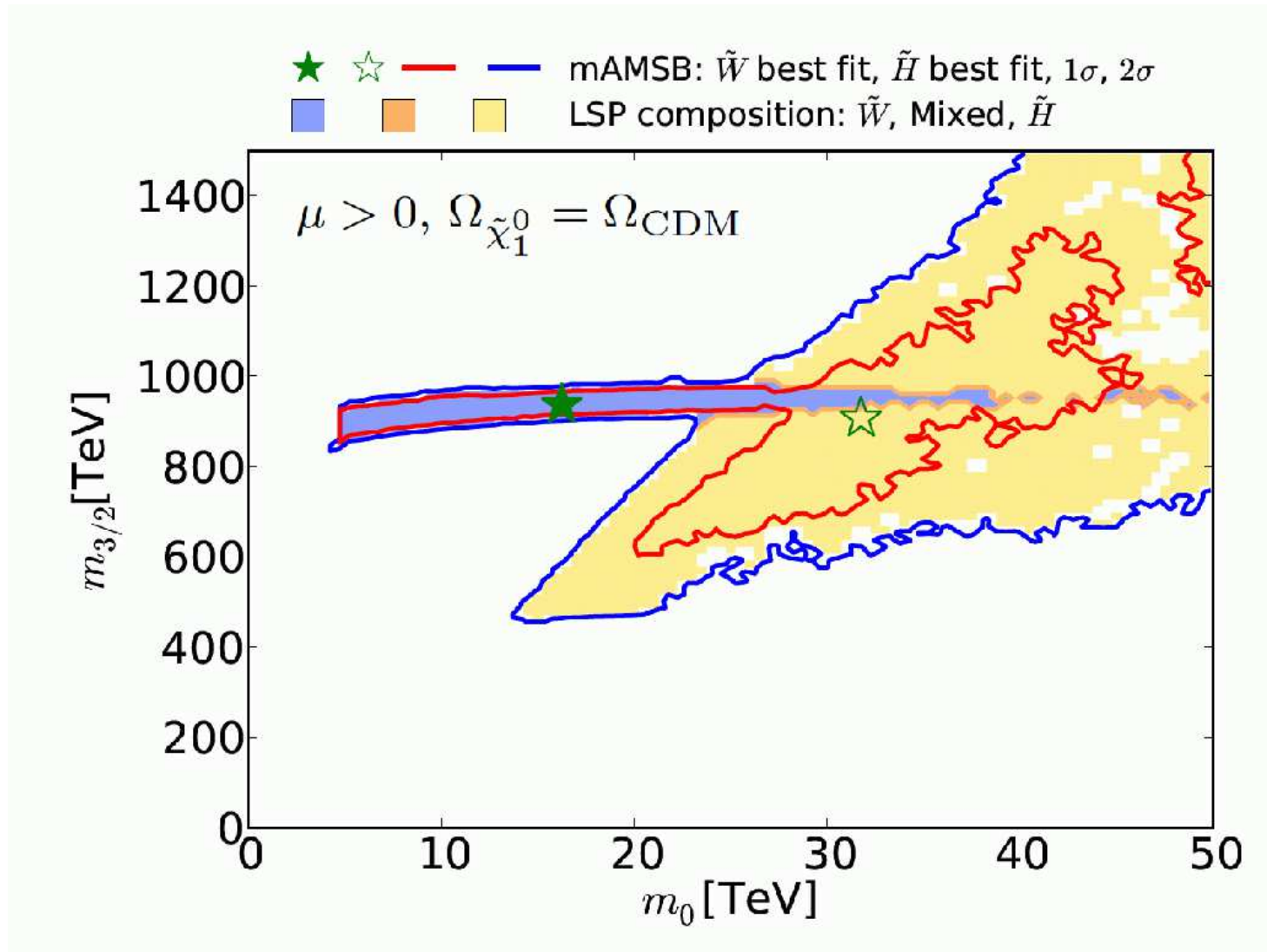


⇒ with Sommerfeld enhancement

⇒ shift to higher $m_{3/2}$

Dark Matter composition:

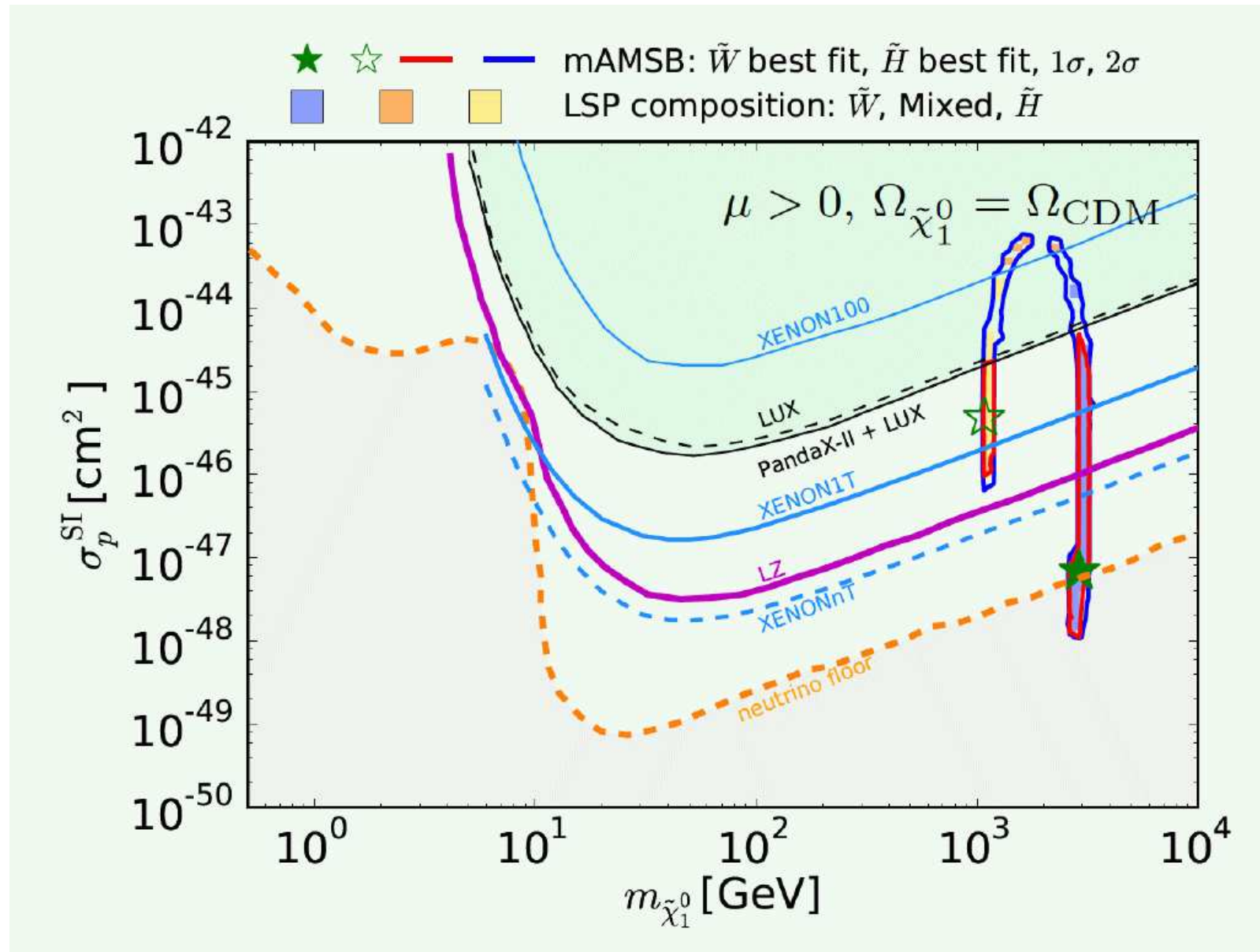
[2016]

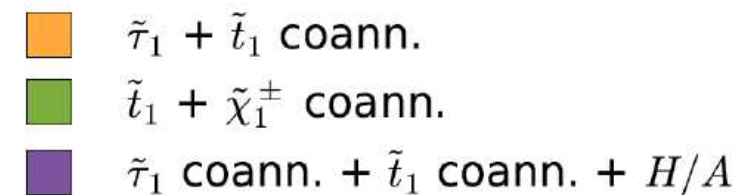
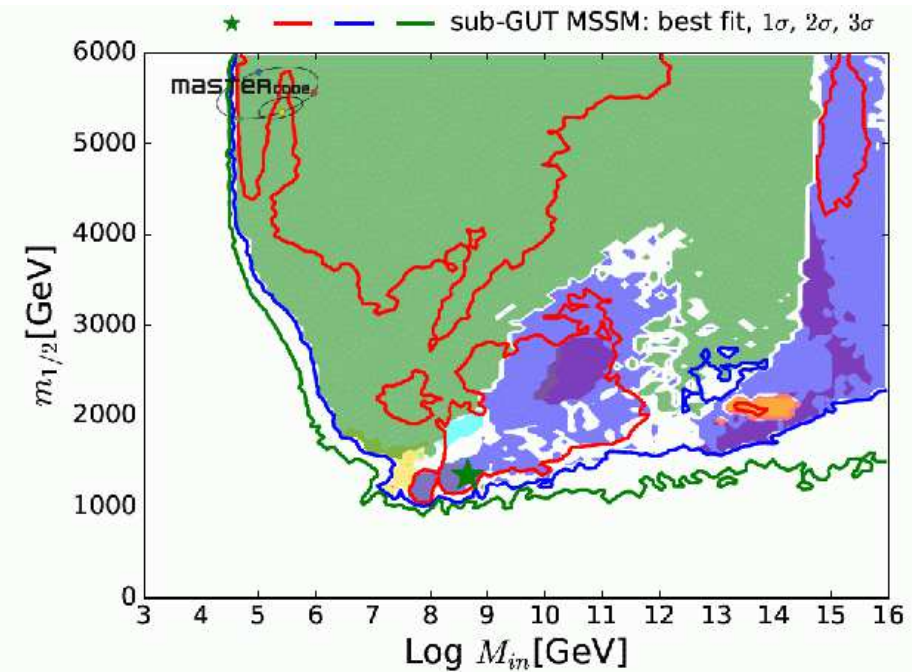
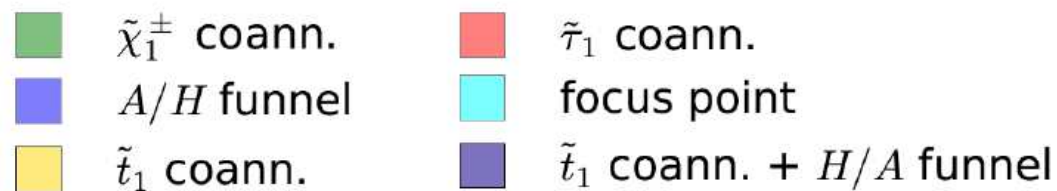
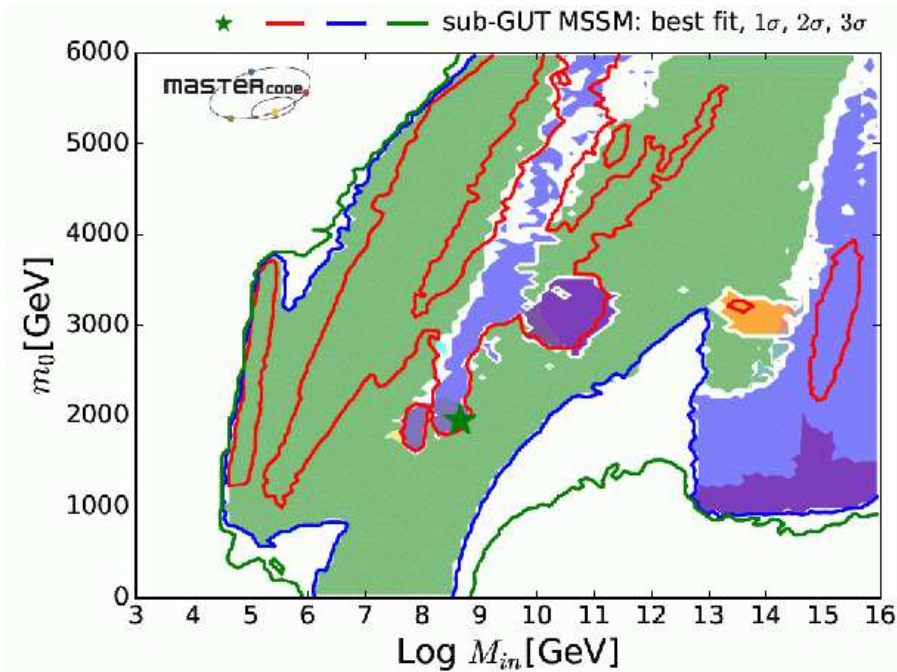


$\Rightarrow m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1 \text{ TeV (wino)}, \sim 1.1 \pm 0.02 \text{ TeV (higgsino)}$

Dark Matter Direct Detection prospects:

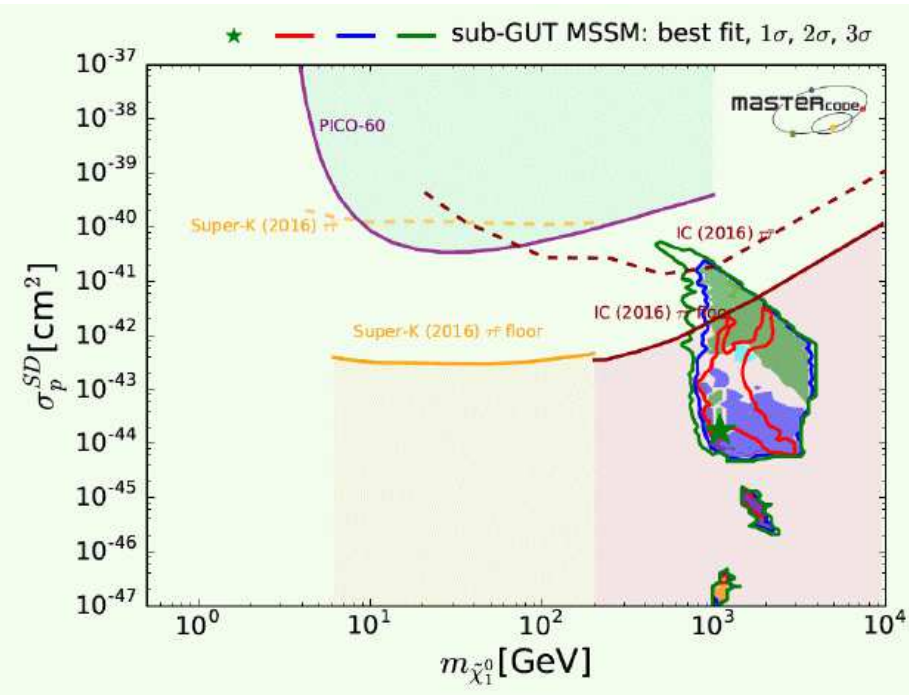
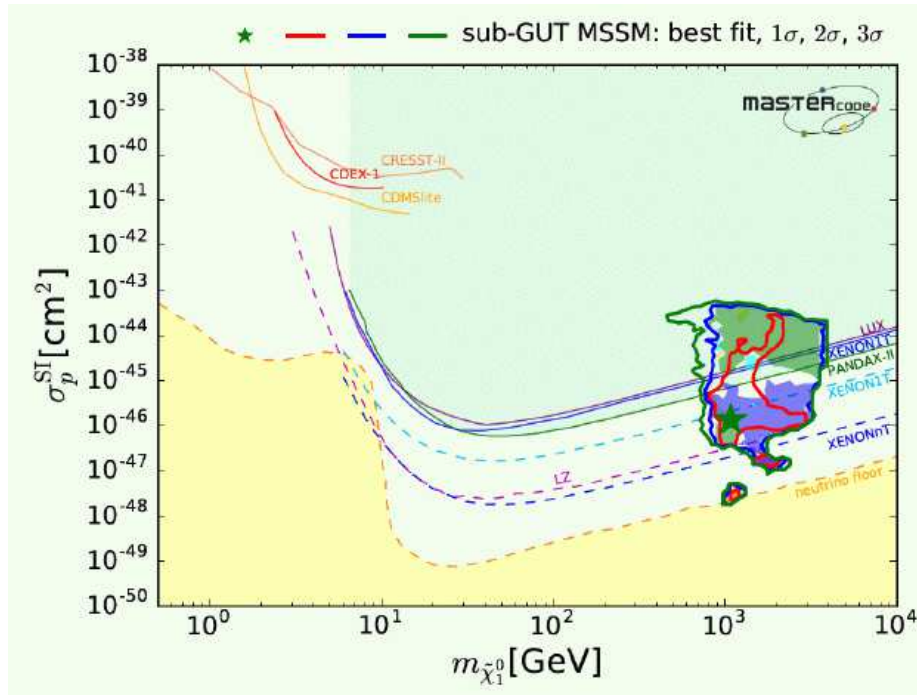
[2016]





⇒ many DM mechanisms possible

⇒ low M_{in} possible/favored



- | | | |
|--|---|--|
| ■ $\tilde{\chi}_1^\pm$ coann. | ■ $\tilde{\tau}_1$ coann. | ■ $\tilde{\tau}_1 + \tilde{t}_1$ coann. |
| ■ A/H funnel | ■ focus point | ■ $\tilde{t}_1 + \tilde{\chi}_1^\pm$ coann. |
| ■ \tilde{t}_1 coann. | ■ \tilde{t}_1 coann. + H/A funnel | ■ $\tilde{\tau}_1$ coann. + \tilde{t}_1 coann. + H/A |

σ_p^{SI} : good prospects, all above the neutrino floor

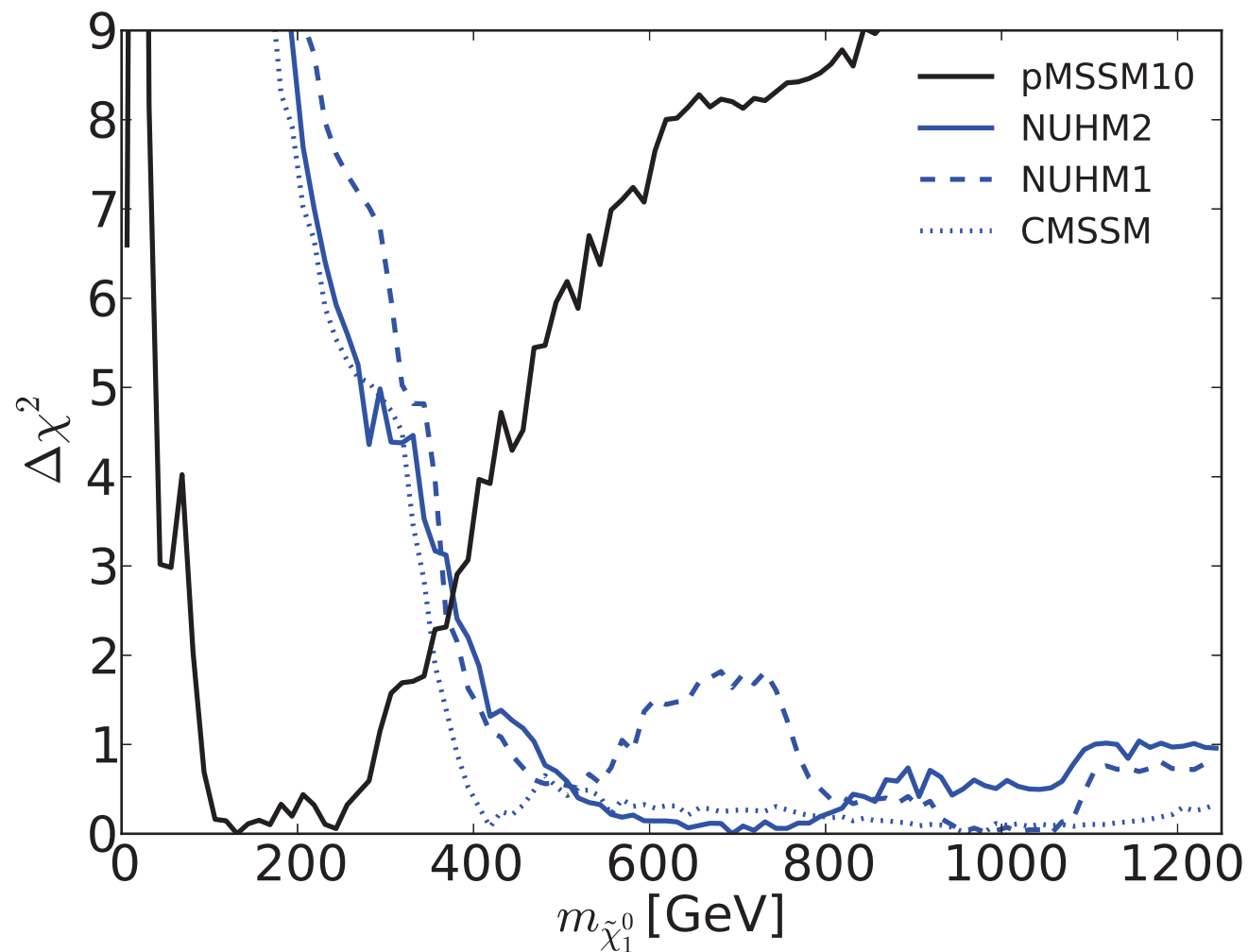
σ_p^{SD} : unclear prospects, best-fit regions below the neutrino floor

3. Results in the pMSSM11



DM mass: pMSSM10 vs. GUT based models prediction:

[2015]

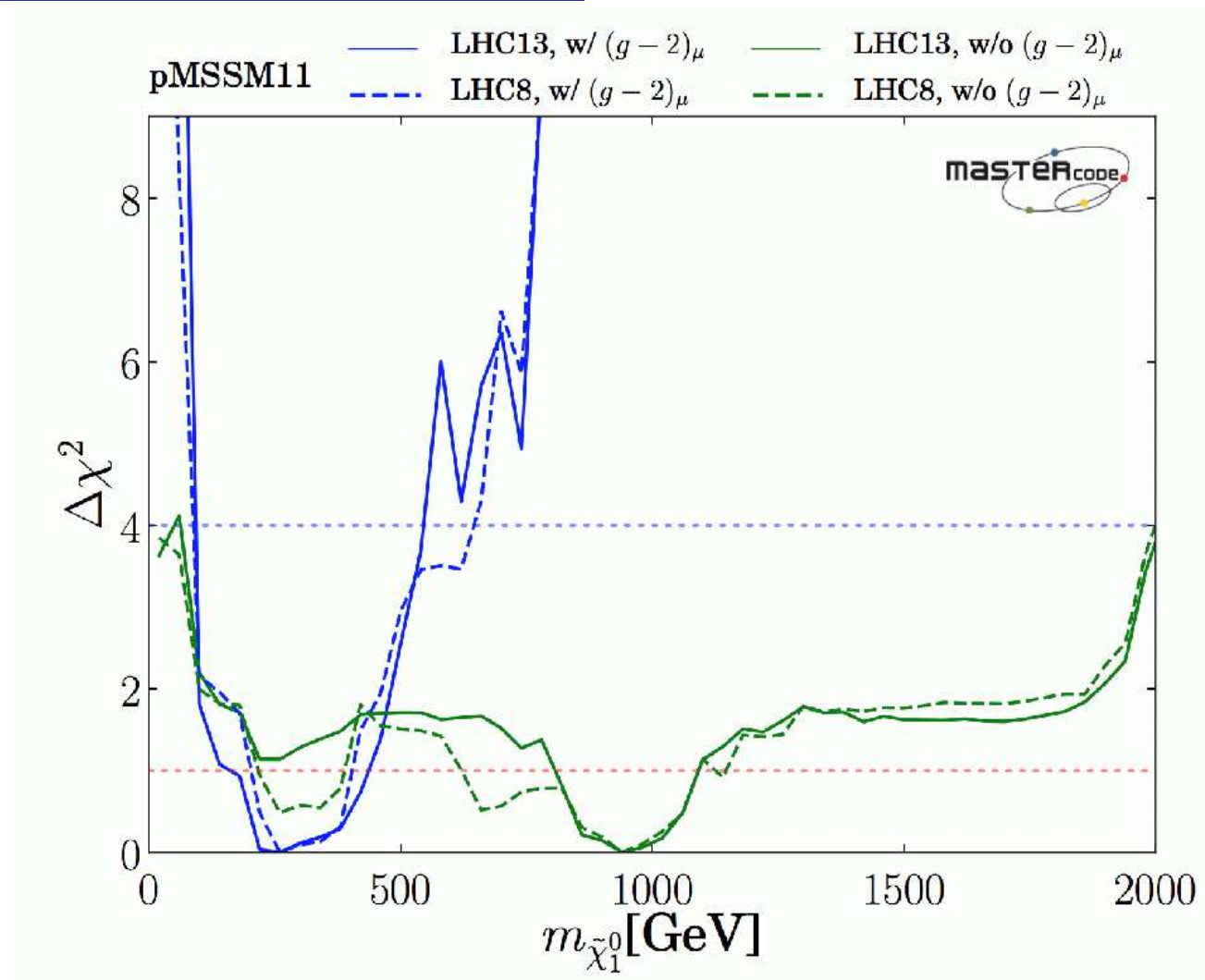


\Rightarrow pMSSM10 predicts much lower DM mass than GUT-based models

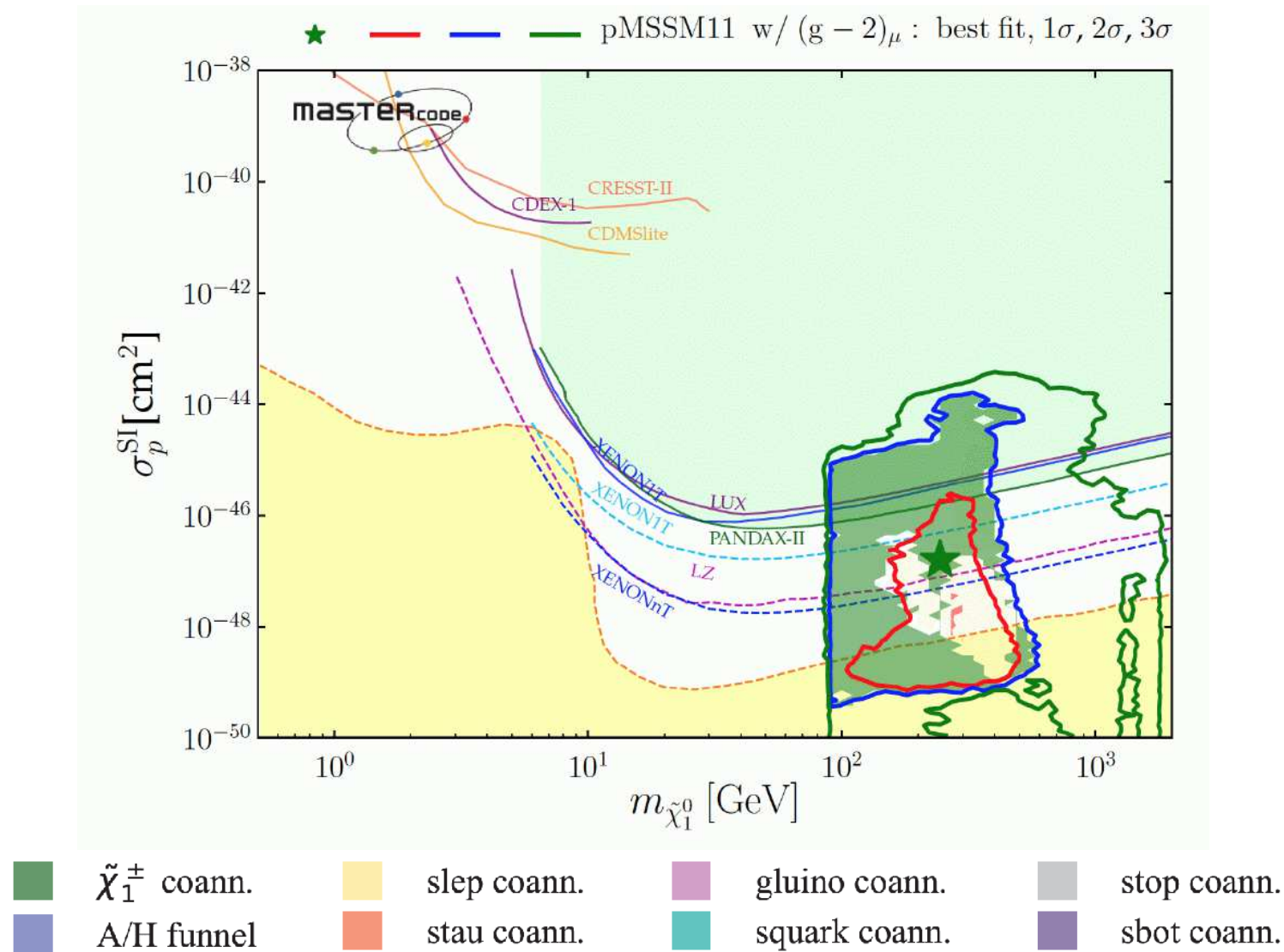
3. Results in the pMSSM11

[2017]

DM mass: similar in the pMSSM11:

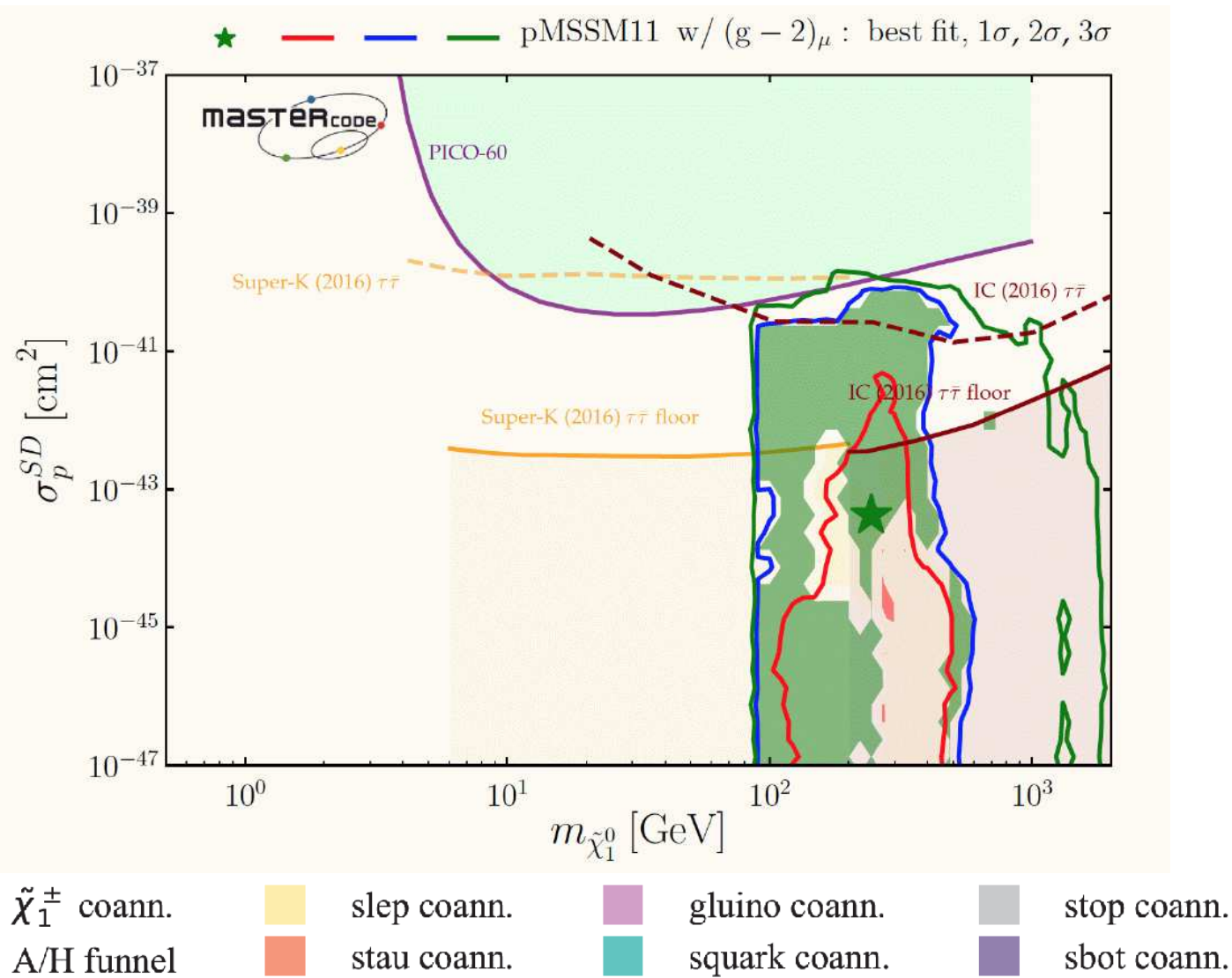


⇒ pMSSM11 predicts much lower DM mass than GUT-based models



⇒ best-fit point covered by future experiments

⇒ but very low cross sections possible at 1σ , below neutrino floor



⇒ slim prospects for future experiments

⇒ large regions allowed at 1 σ , below neutrino floor

4. Results in non-SUSY models

⇒ SM + Dirac DM + Leptophobic spin-1 mediator

Lagrangian according to LHC-DM-WG recommendation:

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Scenarios

- Leptophobic, $g_{l_{i,j}}^V = g_{l_{i,j}}^A = 0$ (no constraints from dilepton searches).
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pure vector.

2. $g_{X_D}^V = 0$ $g_{X_D}^A \equiv g_{DM}$
 $g_{u/d}^V = 0$ $g_{u/d}^A = g_{SM},$

pure axial-vector.

[taken from E. Bagnaschi]

MasterCode set-up :

- Frequentist fitting framework written in Python/Cython and C++
- Multinest algorithm is used to sample the parameter space
- udocker used for deployment

Scan ranges:

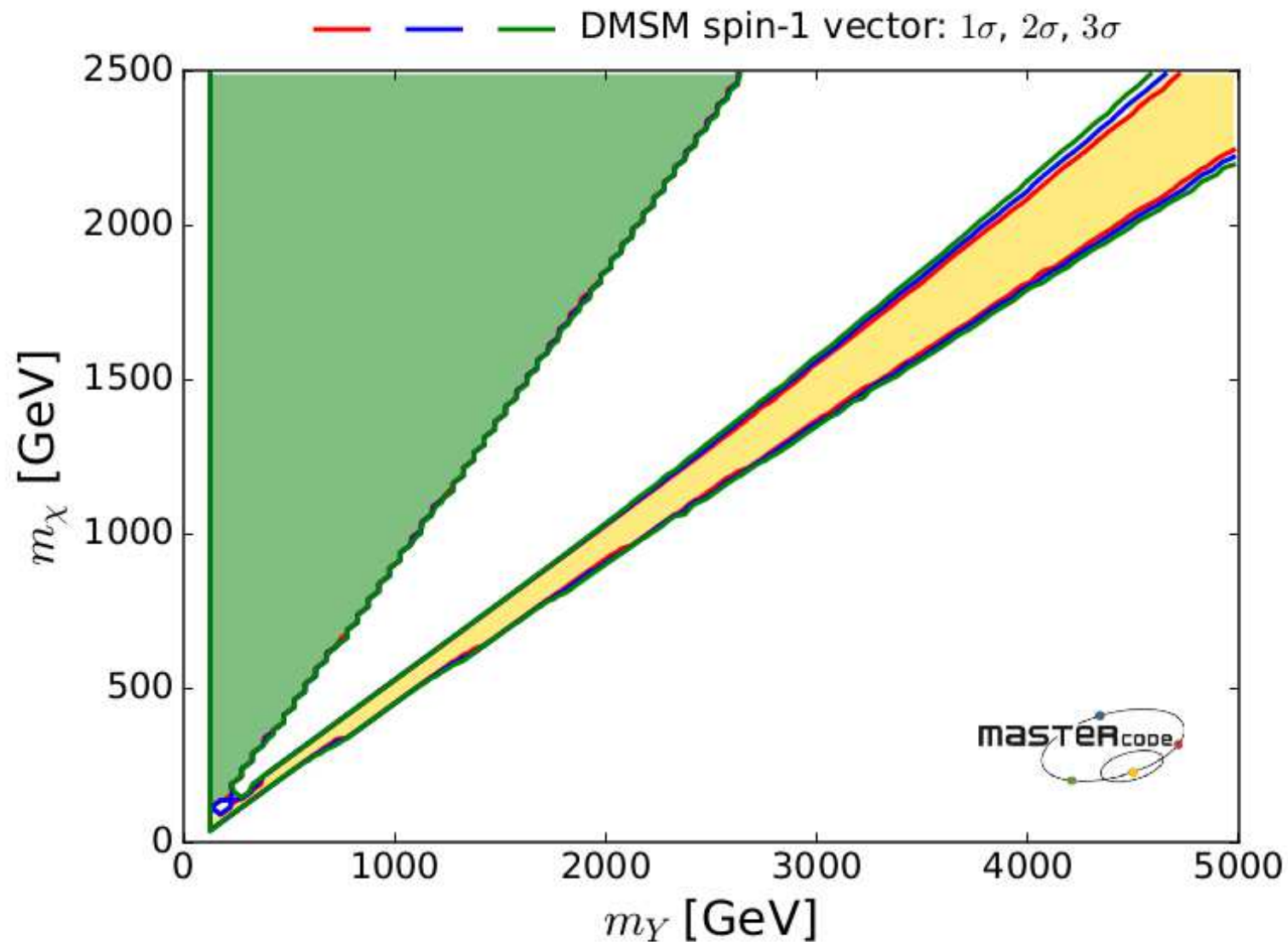
Parameter	Range	# of Segments
m_Y (mediator)	(0.1, 5) TeV	10
m_χ (DM)	(0, 2.5) TeV	8
g_{SM}	$(10^{-6}, \sqrt{4\pi})$	2
g_{DM}	$(10^{-6}, \sqrt{4\pi})$	2
Total # of segments		320

Constraints:

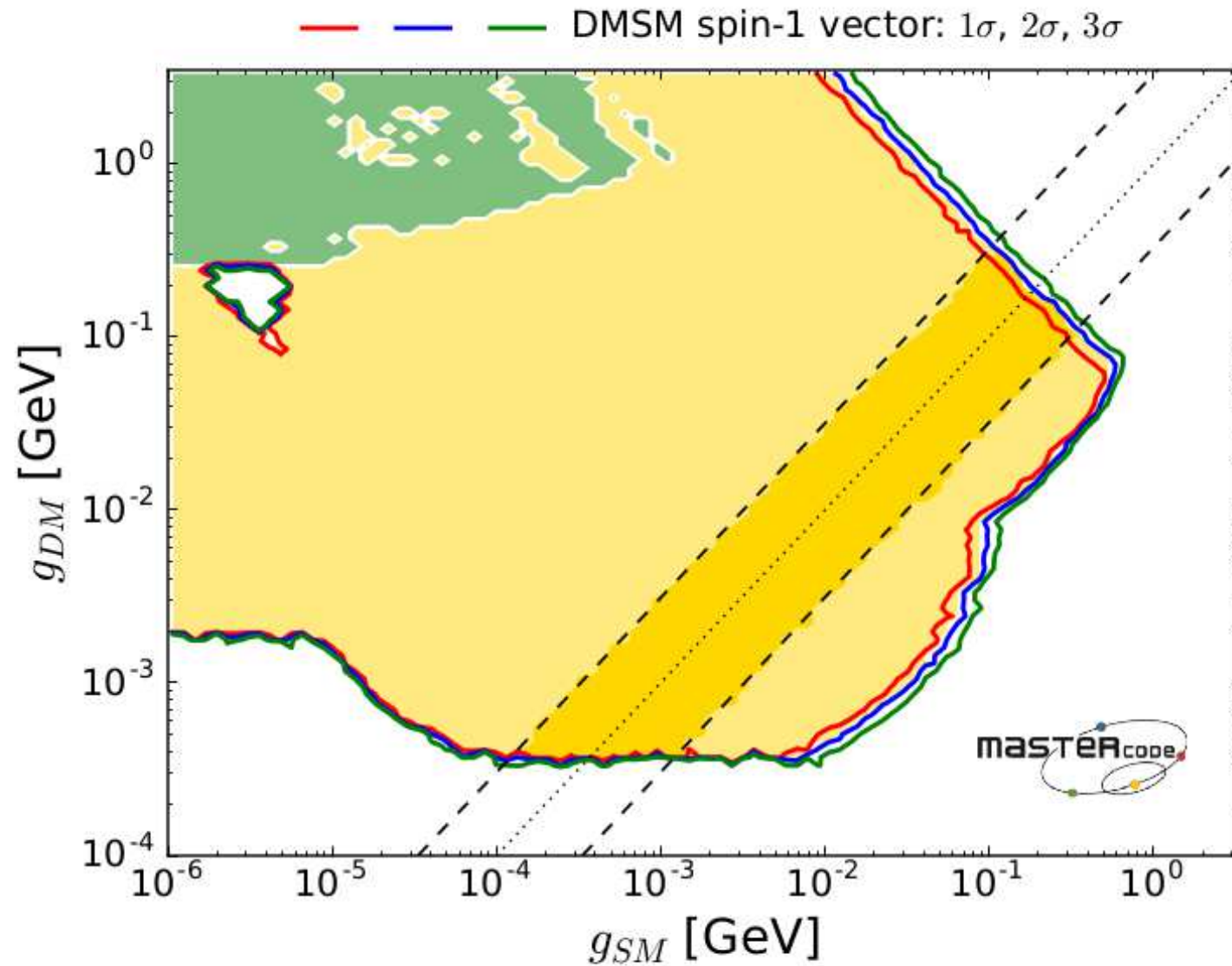
- DM constraints: relic abundance
⇒ full agreement with ATLAS/CMS results/implementations
 - DM constraints: direct detection
⇒ LUX, Xenon1T, PANDAX, PICO60
 - Mono-jet constraints
⇒ MC5 aMC(N)NLO, Fastlim approach
 - di-jet constraints
⇒ MC5 aMC(N)NLO, Fastlim approach
- details in the back-up

General Results

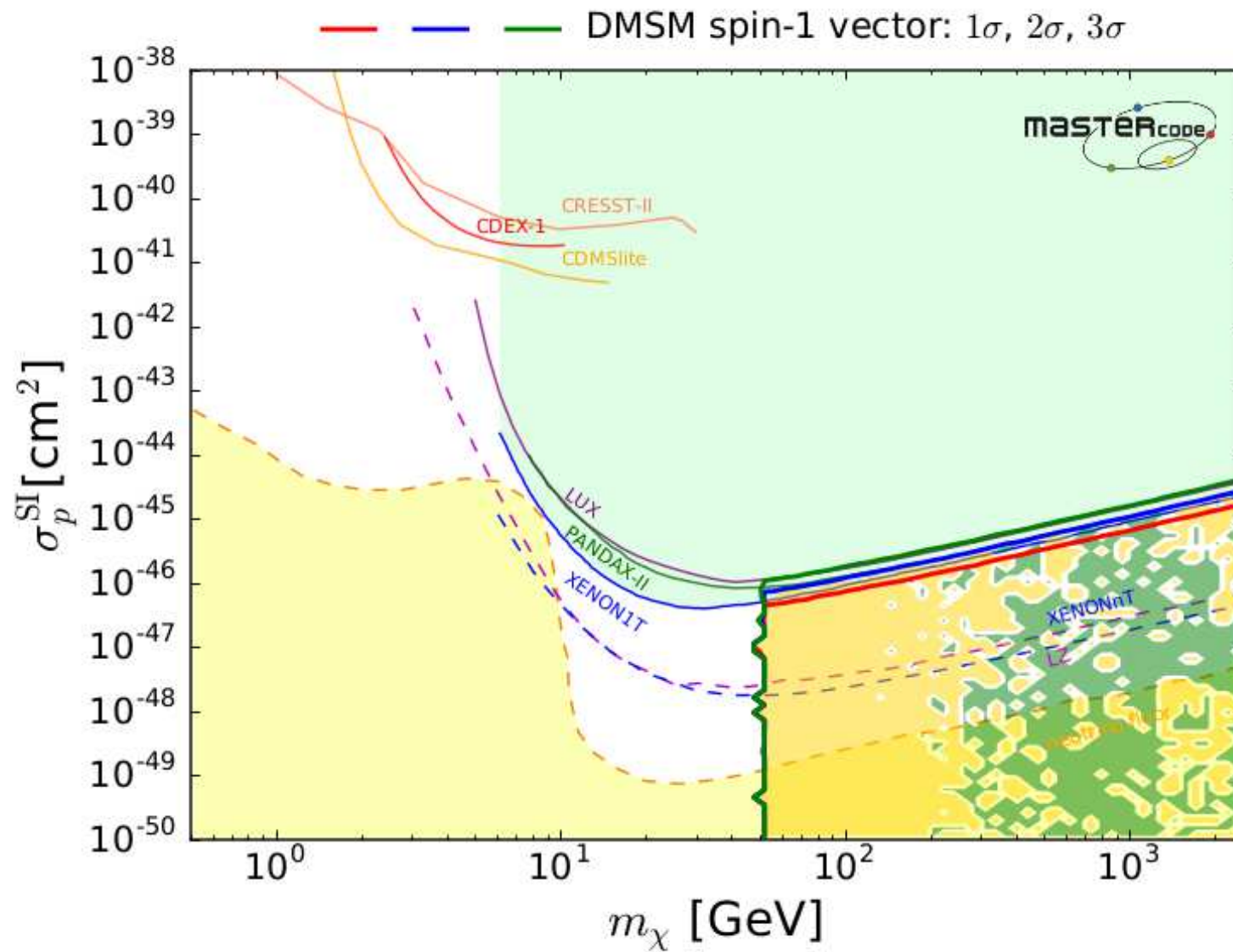
- Results for vector mediator model
- Results for axial-vector mediator model → back-up
- No restrictions on couplings or masses
- Color coding:
 - green: annihilation via t -channel χ exchange
into pairs of mediator particles Y that subsequently decay
into SM particles
 - yellow: rapid annihilation directly into SM particles
via the s -channel Y resonance



⇒ clear separation between s - and t -channel



⇒ large ranges allowed, t -channel only for $g_{DM} \gg g_{SM}$



⇒ mixed prospects, both for s - and t -channel case

Towards UV completions

So far no UV completion considered!

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In any UV completion the spin-one boson could be expected to have comparable couplings to SM and DM particles, modulo possible group-theoretical factors and mixing angles!

$$g_{\text{DM}}/g_{\text{SM}} = \mathcal{O}(1)$$

Towards UV completions

So far no UV completion considered!

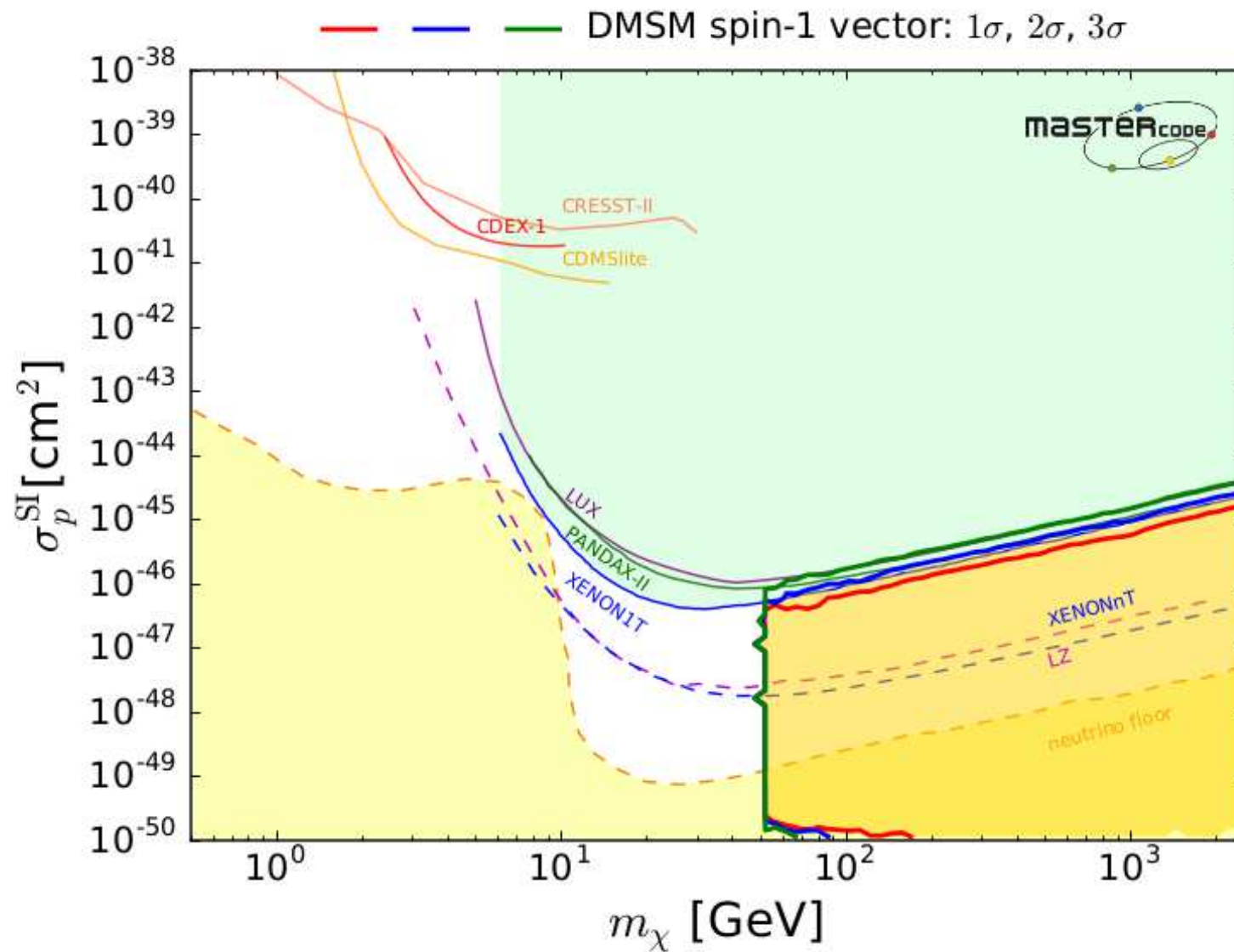
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$$g_{\text{DM}}/g_{\text{SM}} = \mathcal{O}(1)$$

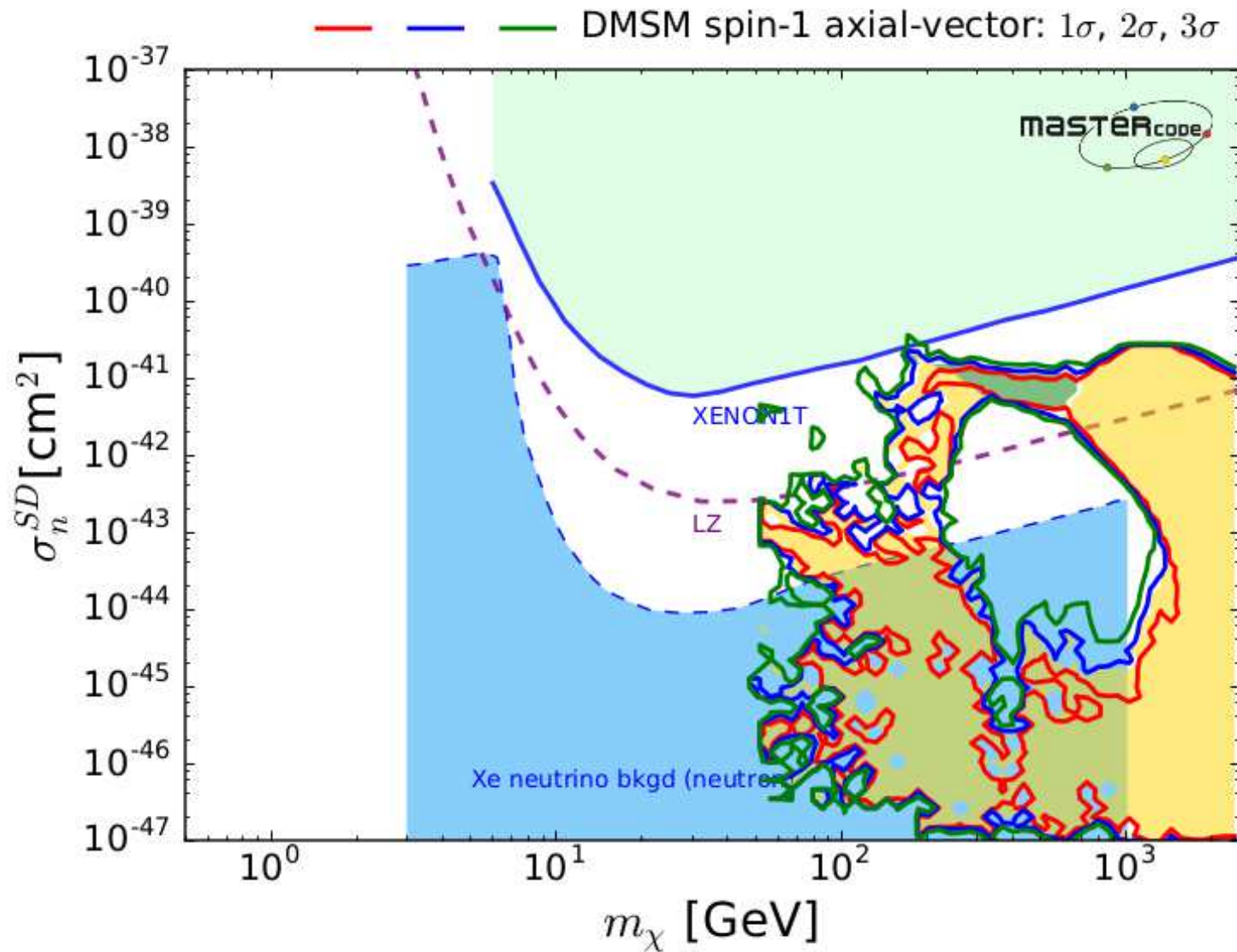
$$1/3 < g_{\text{DM}}/g_{\text{SM}} < 3$$

⇒ dark yellow regions

⇒ *s*-channel favored!



⇒ mixed prospects for discovery



\Rightarrow t -channel can fully be probed, s -channel only partially

5. Conclusinos

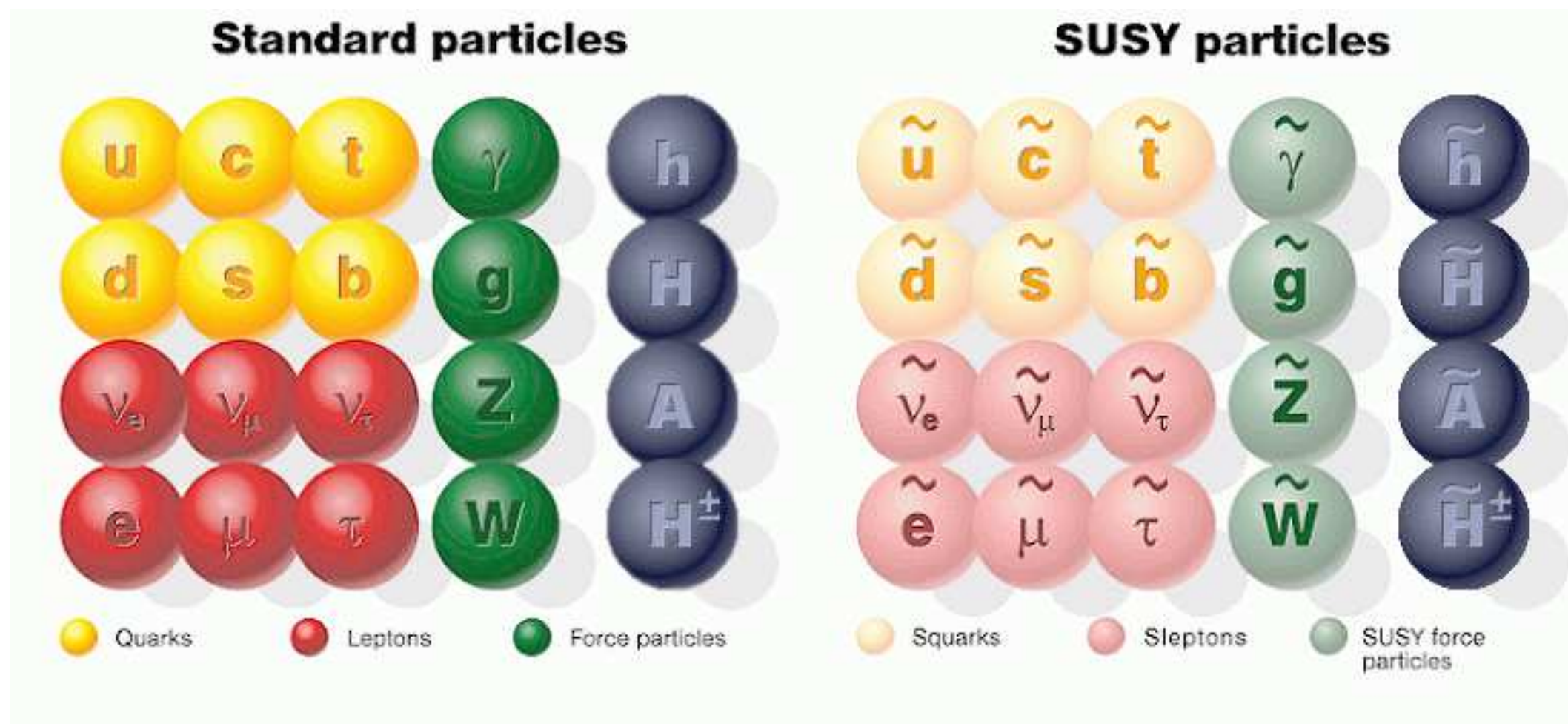
- SUSY constrained: CMSSM, NUHM, SU(5), mAMSB, sub-GUT
SUSY general: pMSSM11, ...
- Our tool: MasterCode: combination of all relevant data!
- CMSSM, NUHM1, NUHM2:
 - $m_{\tilde{\chi}_1^0} \gtrsim 400$ GeV neutrino floor
 - best-fit regions (mostly) above/at/below
- SU(5):
 - stau co-ann., but also $\tilde{u}_R/\tilde{c}_R/\tilde{\nu}_\tau$ co-ann. possible
 - $m_{\tilde{\chi}_1^0}$ as in CMSSM, NUHM1, NUHM2
- mAMSB:
 - $m_{\tilde{\chi}_1^0} \sim 2.9 \pm 0.1$ TeV (wino), $\sim 1.1 \pm 0.02$ TeV (higgsino)
 - DD: wino at neutrino floor, higgsino tested by next round
- pMSSM11:
 - $m_{\tilde{\chi}_1^0} \lesssim 500$ GeV; important: chargino co-annihilation
 - σ_p^{SI} partially cov. at future exp., σ_p^{SD} below neutrino floor
- SM + Dirac DM + Leptophobic spin-1 mediator
⇒ MasterCode approach: full fit of the model (no simplifying ass.)
Vector mediator: s - and t -channel separated, mixed prospects for DD
Axialvector: s - and t -channel continuous, mixed prospects for DD
UV-completions: $1/3 < g_{\text{SM}}/g_{\text{DM}} < 3 \Rightarrow s$ -channel preferred
⇒ prospects for DD not improved



Further Questions?

The Minimal Supersymmetric Standard Model (MSSM)

Superpartners for Standard Model particles



Problem in the MSSM: more than 100 free parameters

Nobody(?) believes that a model describing nature has so many free parameters!

A. Unconstrained models (MSSM):

agnostic about how SUSY breaking is achieved

no particular SUSY breaking mechanism assumed, parameterization of possible soft SUSY-breaking terms

most general case:

⇒ 105 new parameters: masses, mixing angles, phases

⇒ no model missed (within the MSSM)

⇒ $\mathcal{O}(100)$ parameters difficult to handle

B. Constrained models:

CMSSM, NUHM1, NUHM2, SU(5), mAMSB, sub-GUT, ...:

assumption on the scenario that achieves spontaneous SUSY breaking

⇒ prediction for soft SUSY-breaking terms
in terms of small set of parameters

⇒ easy to handle

GUT based models: 1.) CMSSM (sometimes wrongly called mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

m_0 : universal scalar mass parameter
 $m_{1/2}$: universal gaugino mass parameter
 A_0 : universal trilinear coupling
 $\tan \beta$: ratio of Higgs vacuum expectation values
 $\text{sign}(\mu)$: sign of supersymmetric Higgs parameter

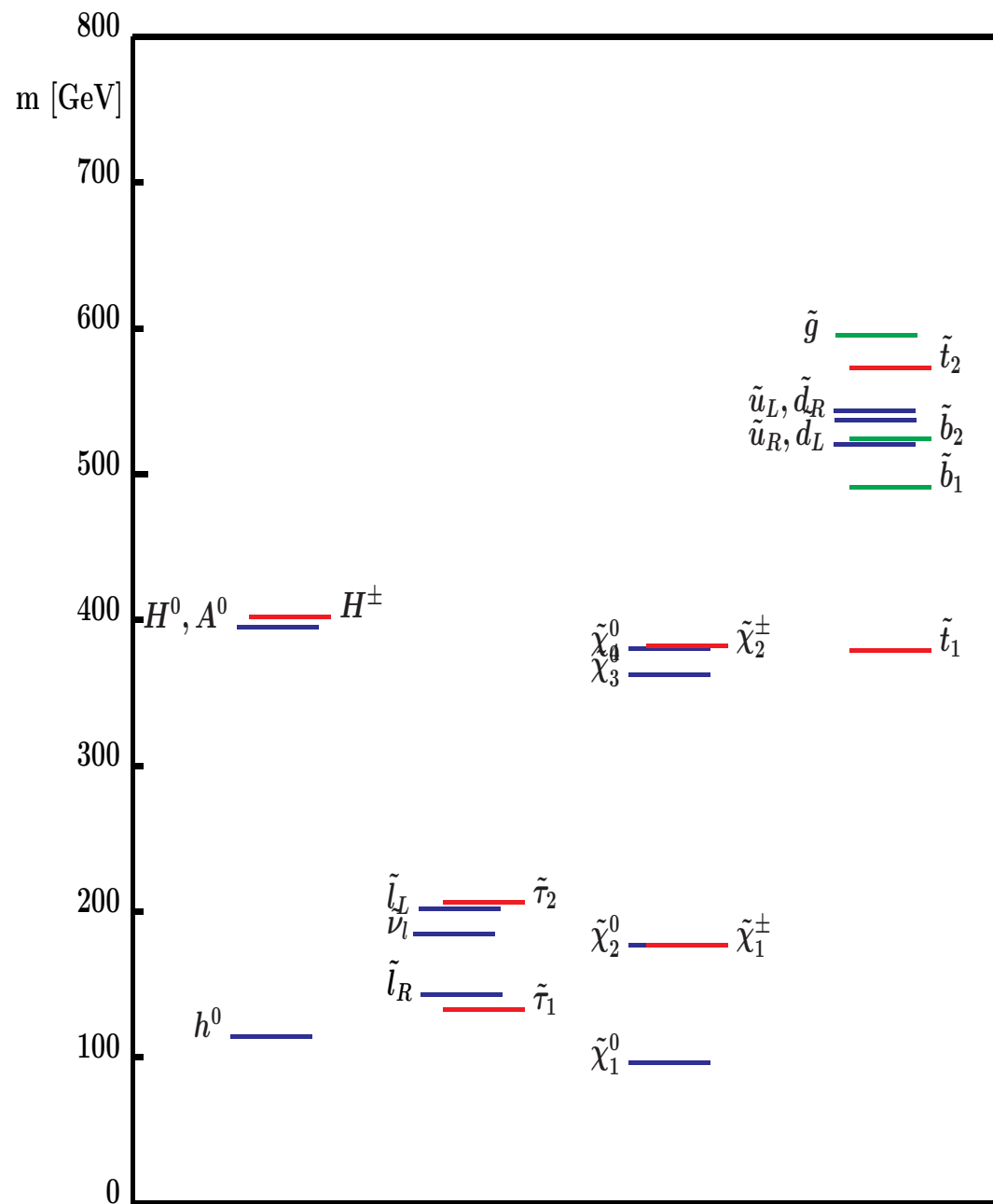
} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

⇒ Lightest SUSY particle (LSP) is the lightest neutralino ⇒ DM!

“Typical” CMSSM scenario
(SPS 1a benchmark scenario):

Strong connection between
all the sectors



GUT based models: 2.) NUHM1: (Non-universal Higgs mass model)

Assumption: no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively M_A as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu \text{ and } M_A$$

GUT based models: 3.) NUHM2: (Non-universal Higgs mass model 2)

Assumption: no unification of scalar Higgs parameter at the GUT scale

⇒ effectively M_A and μ as free parameters at the EW scale

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \mu \text{ and } M_A$$

GUT based models: 4.) SU(5) GUT:

Assumption I:

no unification of scalar Higgs parameter at the GUT scale

(\Rightarrow effectively M_A and μ as free parameters at the EW scale)

Assumption II:

$$(q_L, u_L^c, e_L^c)_i \in \mathbf{10}_i, \quad (\ell_L, d_L^c)_i \in \bar{\mathbf{5}}_i$$

\Rightarrow Scenario characterized by

$$m_5, m_{10}, m_{1/2}, A_0, \tan \beta, m_{H_u}, m_{H_d}$$

GUT based models: 5.) mAMSB:

mAMSB scenario characterized by

$$m_{3/2}, m_0, \tan \beta, \text{sign}(\mu)$$

$m_{3/2} = \langle F \rangle / M_{\text{Planck}}$: overall scale of SUSY particle masses

m_0 : phenomenological parameter: universal scalar mass term introduced in order to keep squares of slepton masses positive

typical feature: very small neutralino–chargino mass difference
 $\Rightarrow \tilde{\chi}_1^\pm \rightarrow \tilde{\chi}_1^0 + \pi^\pm$ with very soft pions

GUT based models: 6.) sub-GUT:

Based on CMSSM with unification at $M_{\text{GUT}} \sim 2 \cdot 10^{16}$ GeV:

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

Unification is assumed at $M_{\text{in}} \leq M_{\text{GUT}}$:

⇒ Scenario characterized by

$$M_{\text{in}}, m_0, m_{1/2}, A_0, \tan \beta, \text{sign } \mu$$

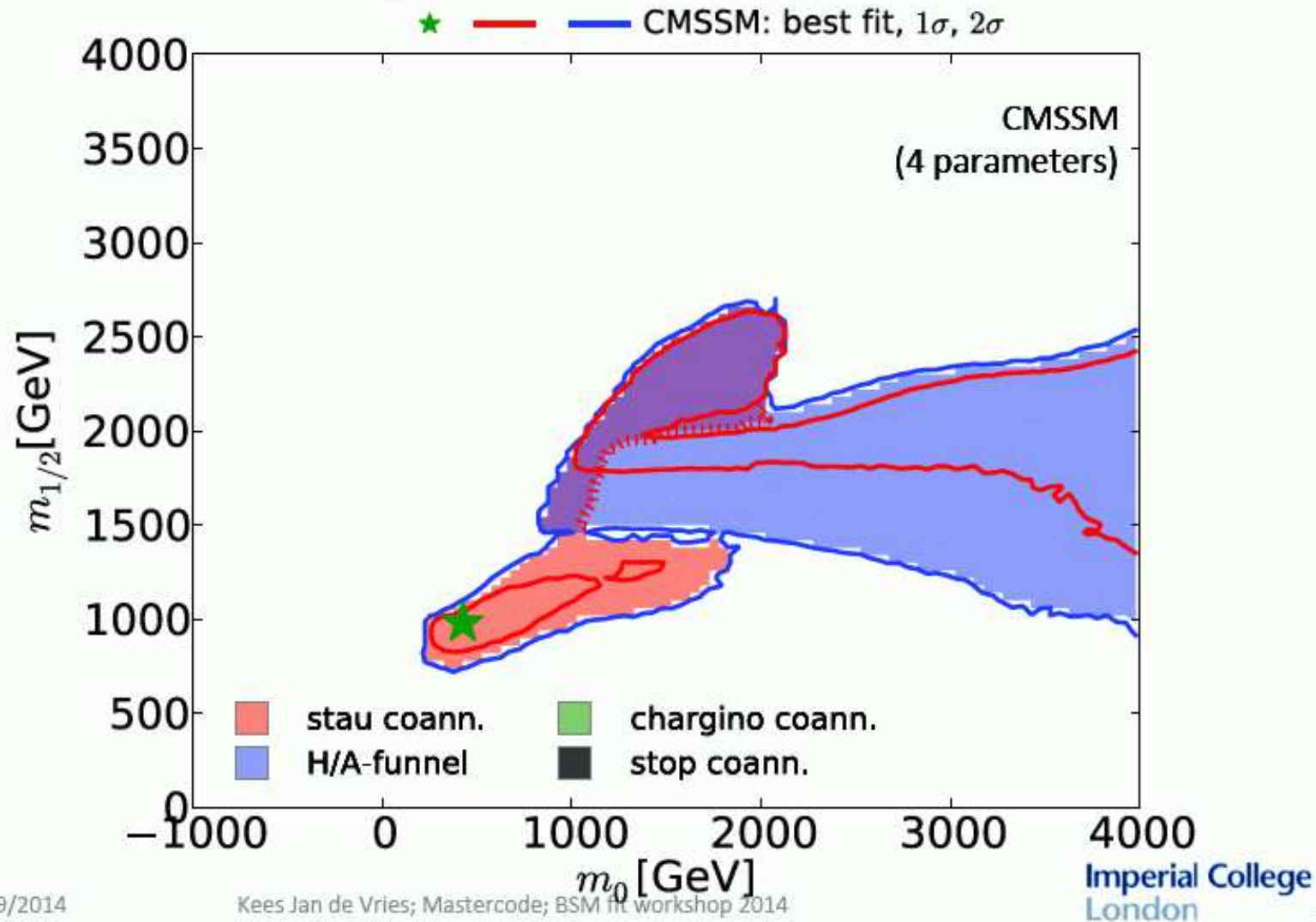
Possible realization in “mirage unification”

warped extra dimensions

...



Mechanisms for relic dark matter density fulfillment in the CMSSM

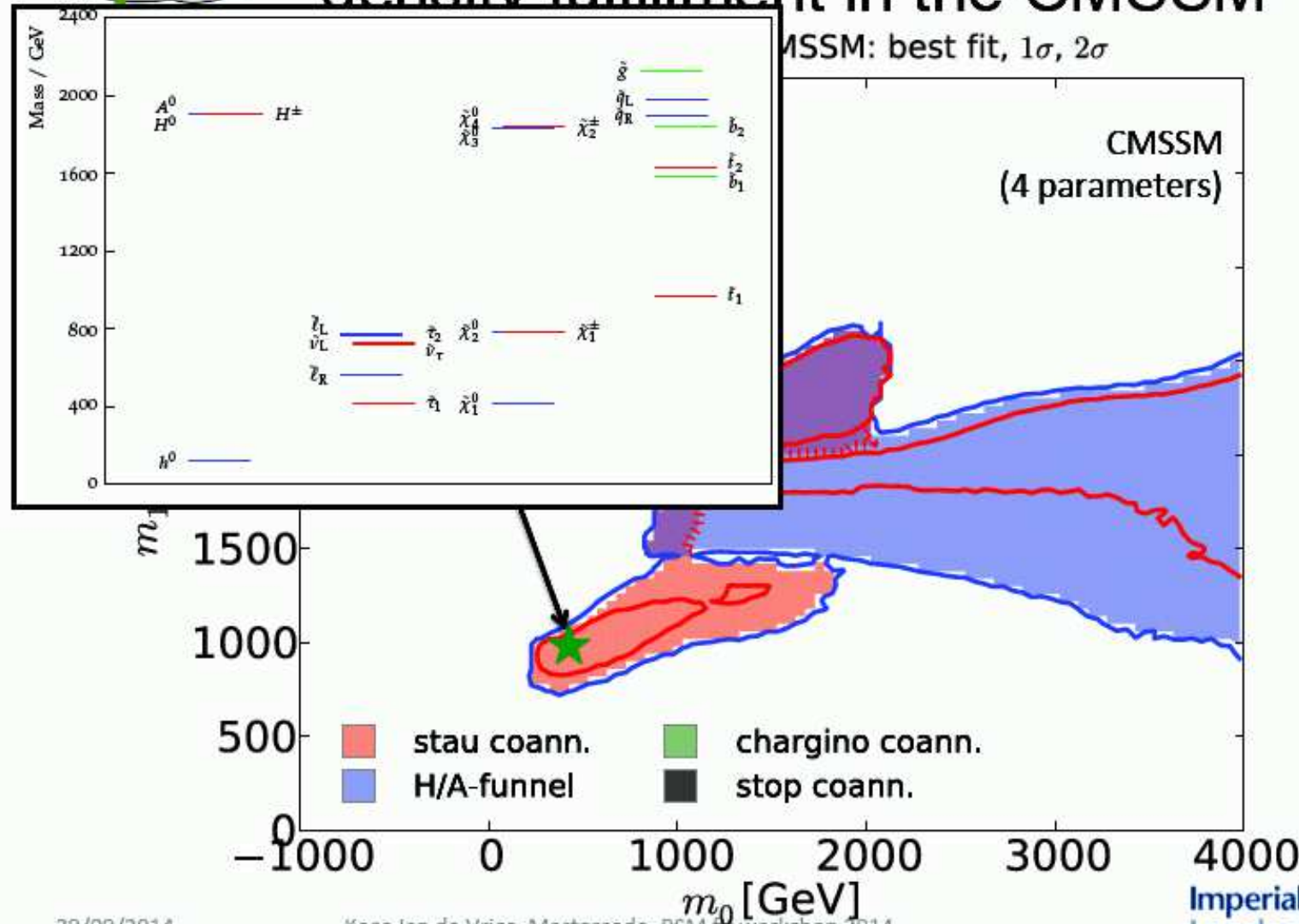


29/09/2014

Kees Jan de Vries; Mastercode; BSM fit workshop 2014



Mechanisms for relic dark matter density fulfillment in the CMSSM



29/09/2014

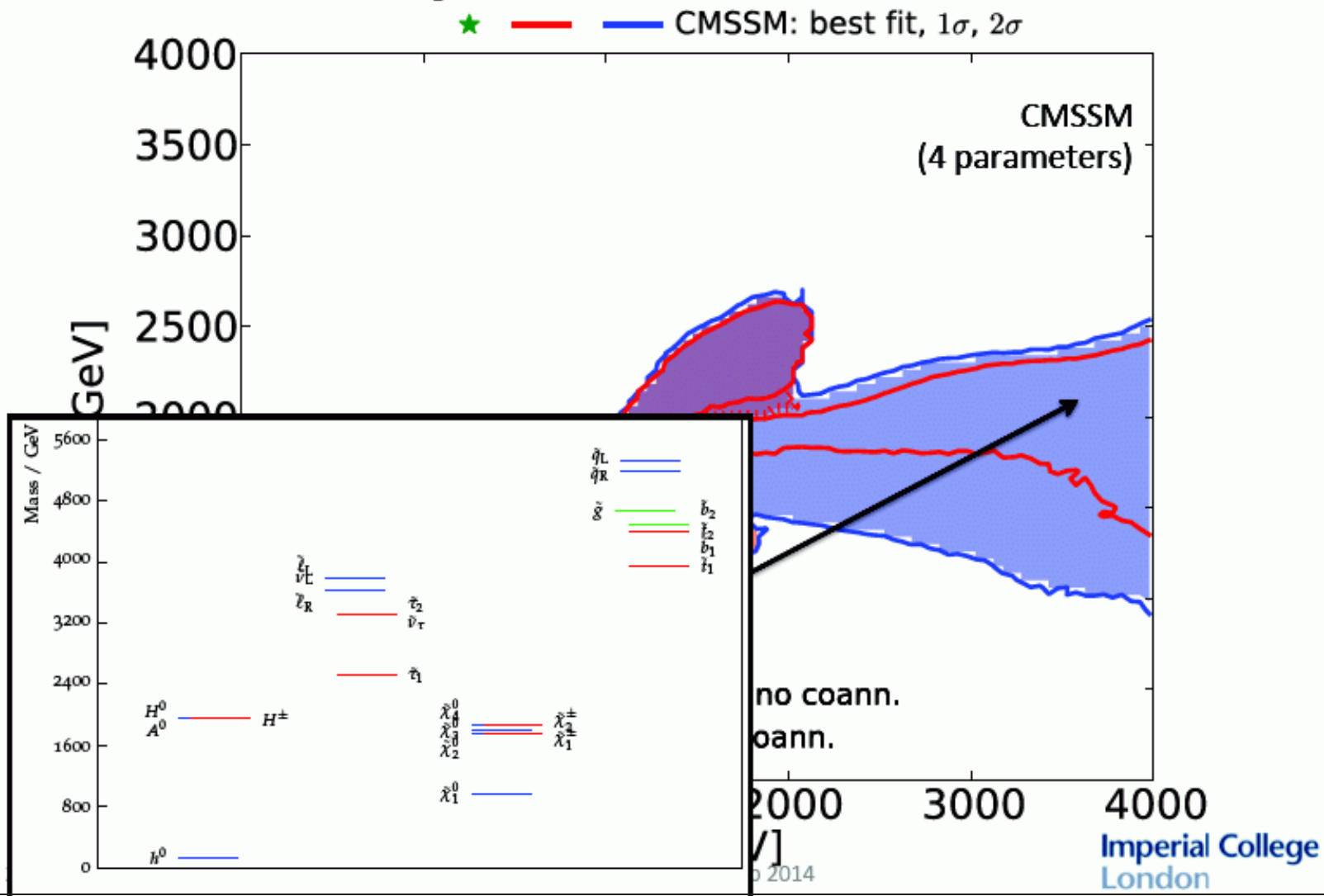
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8

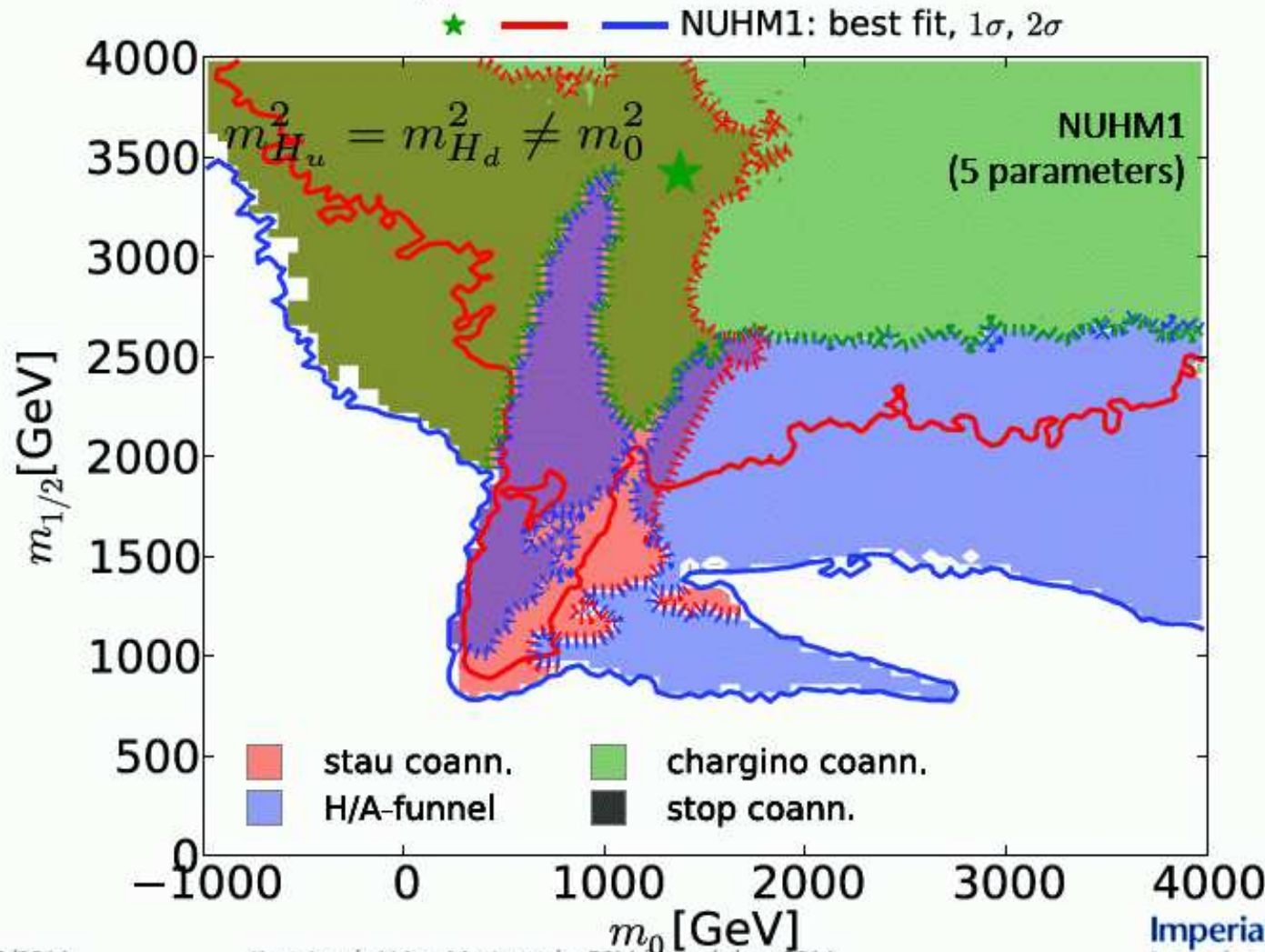


Mechanisms for relic dark matter density fulfillment in the CMSSM





Mechanisms for relic dark matter density fulfillment in the NUHM1



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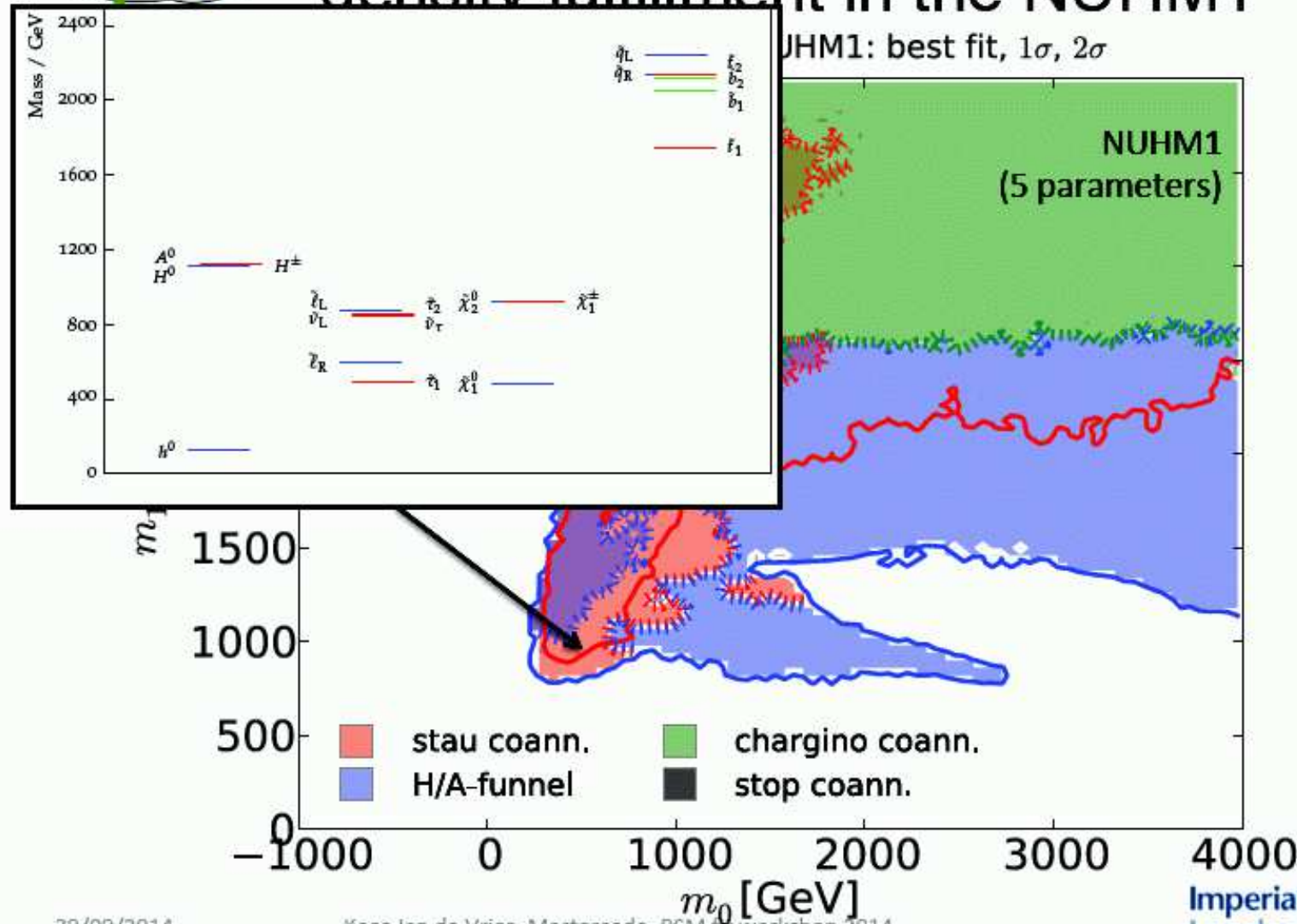
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10



Mechanisms for relic dark matter density fulfillment in the NUHM1



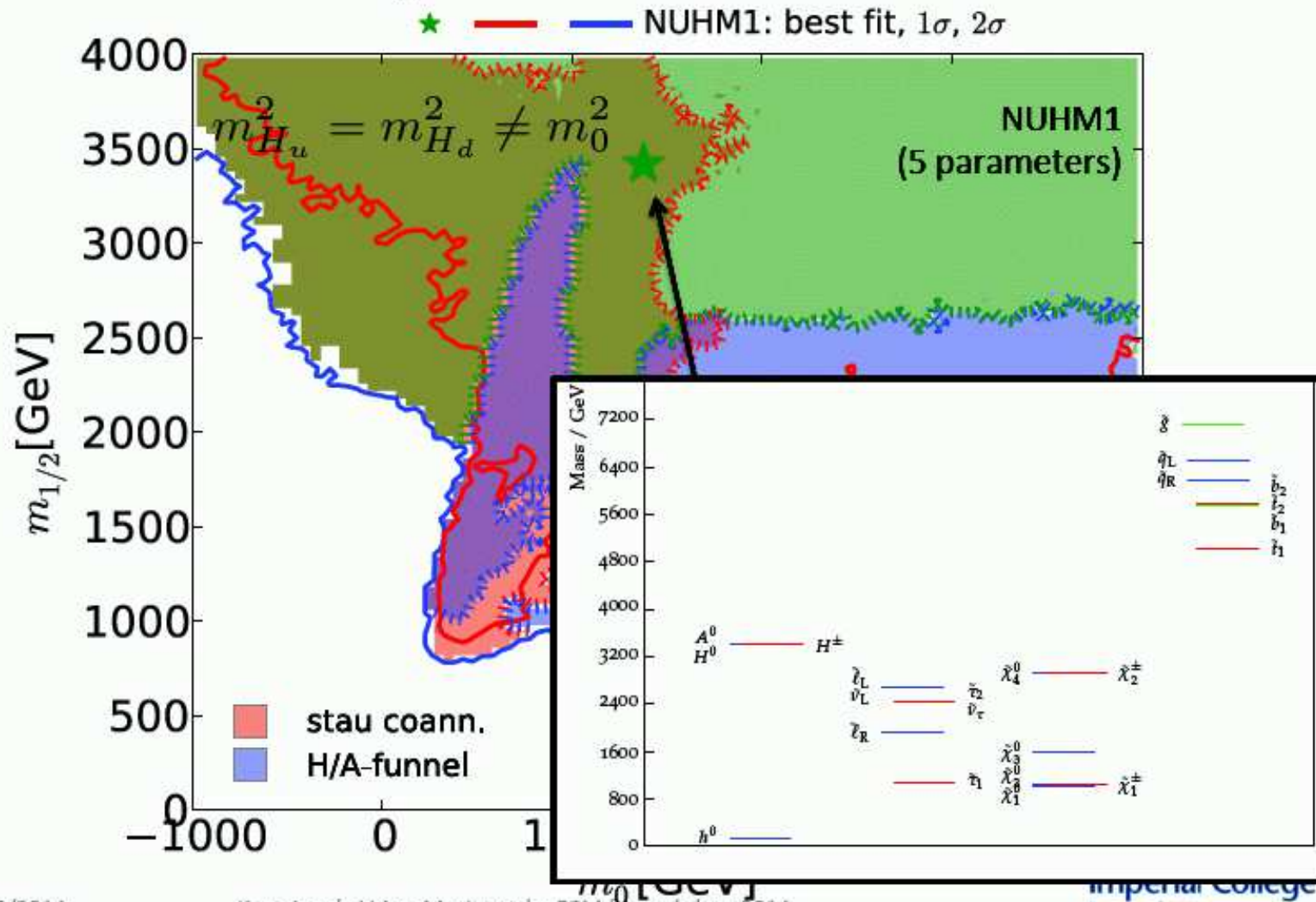
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11



Mechanisms for relic dark matter density fulfillment in the NUHM1



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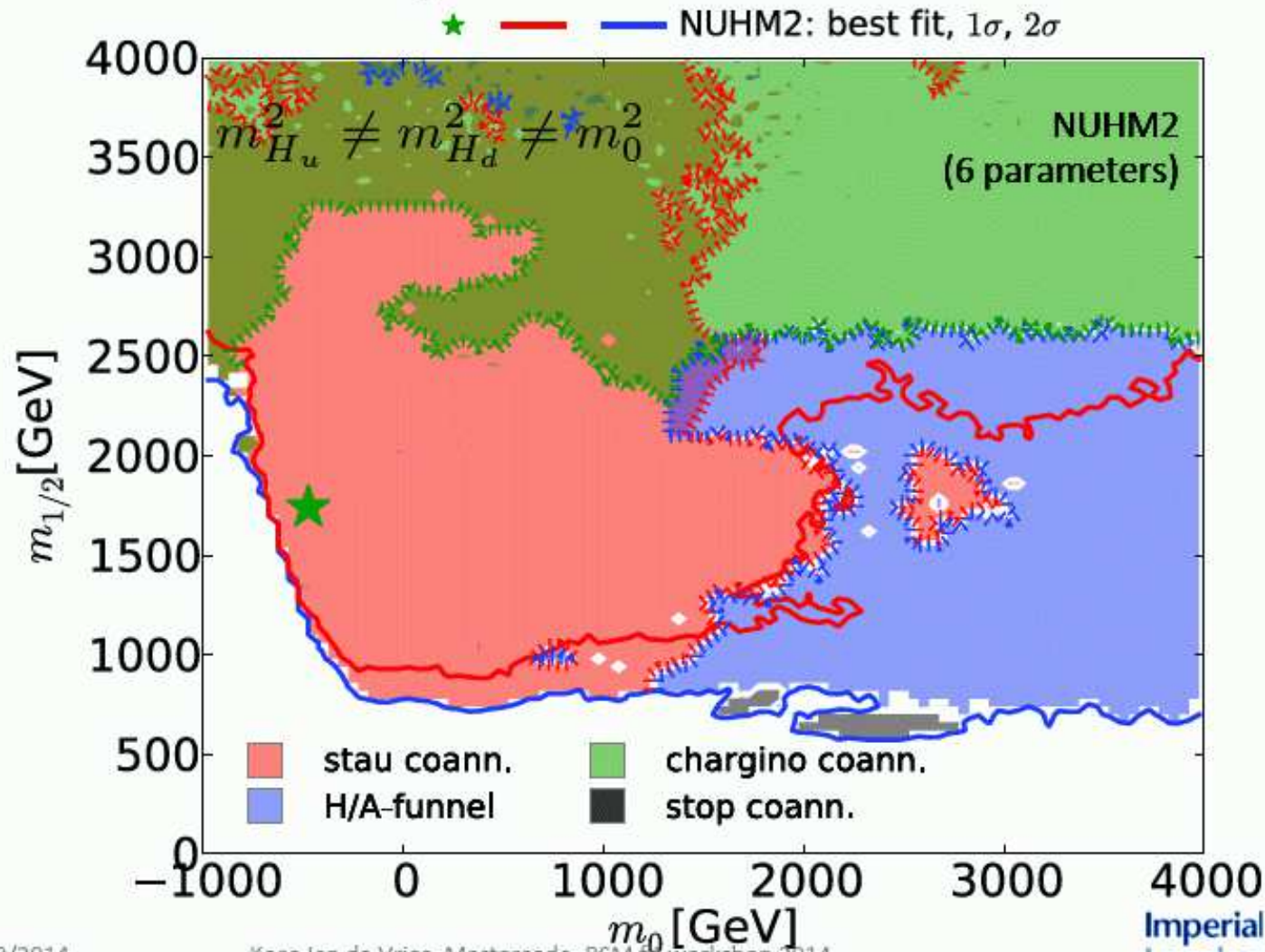
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12



Mechanisms for relic dark matter density fulfillment in the NUHM2

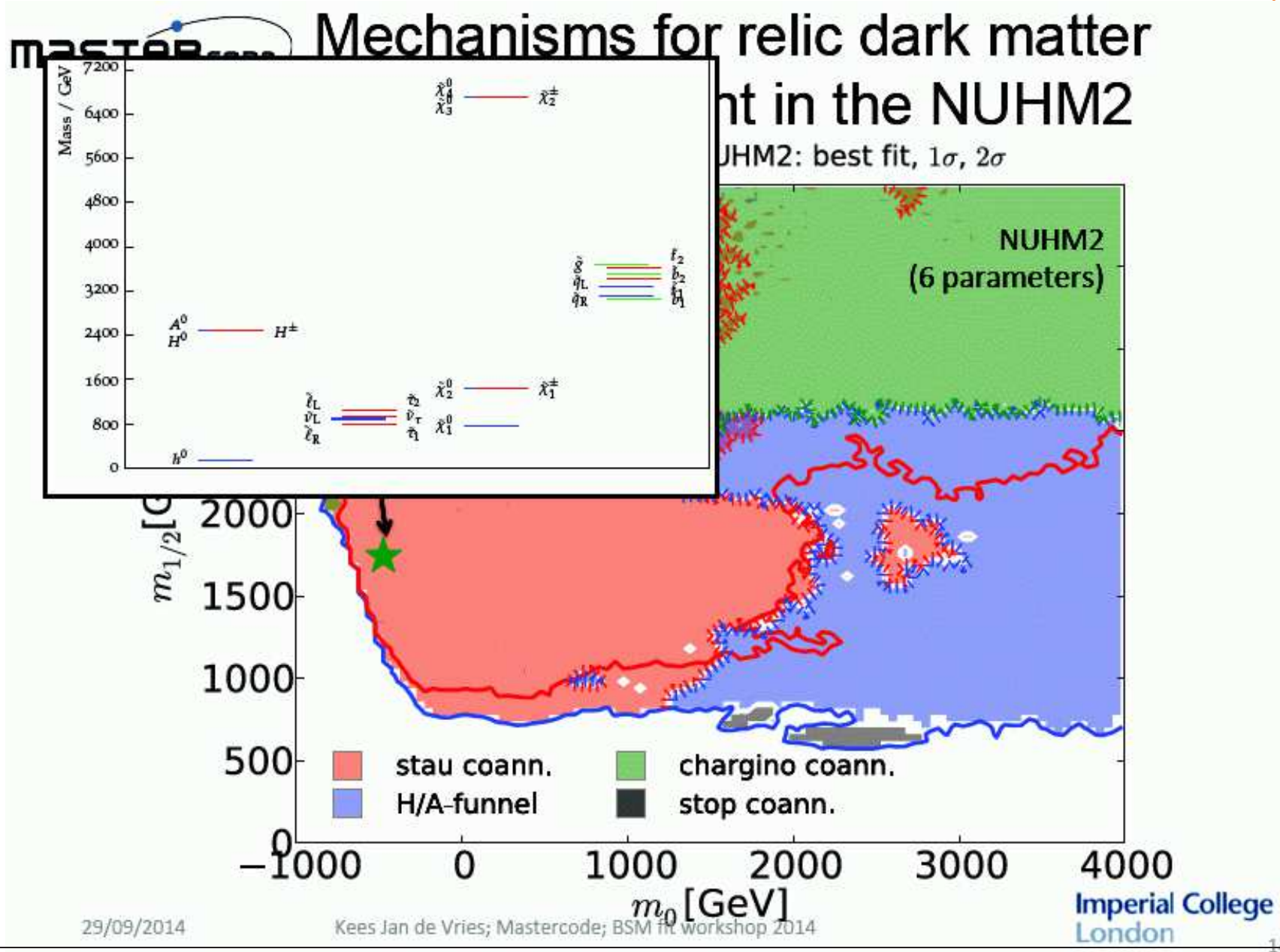


29/09/2014

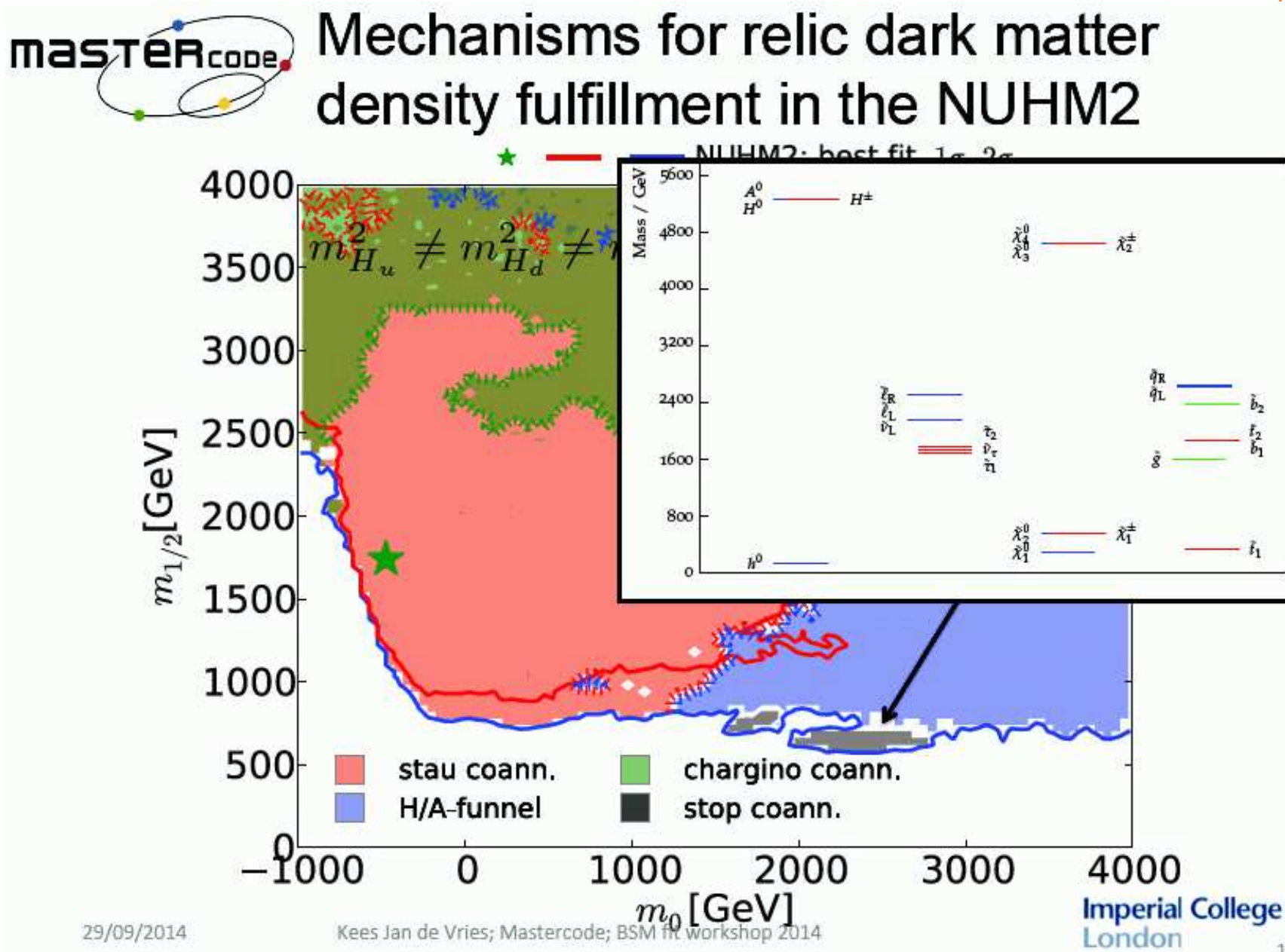
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13



「2014」



5. Conclusinos

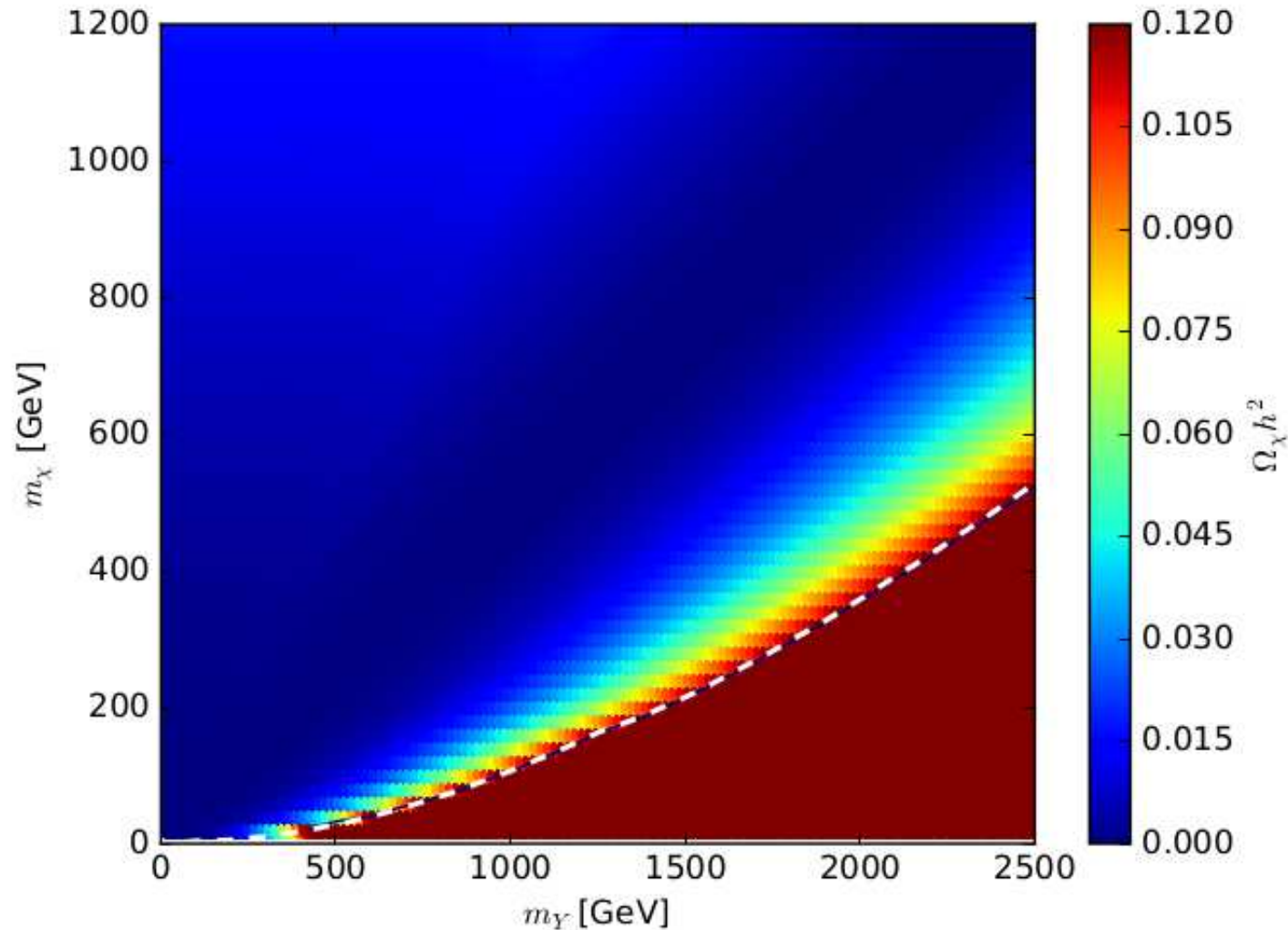
- **SUSY** is (still) the best-motivated BSM scenario
 - constrained models: **CMSSM**, **NUHM**, **SU(5)**, **mAMSB**, **sub-GUT**
 - general models: **pMSSM11**, ...
- Our tool: **MasterCode**: combination of **LHC** searches, Higgs measurements, **EWPO**, **BPO**, **CDM** $\Rightarrow \chi^2$ evaluation
- Results wrt. neutrino floors:

Model	Min. χ^2/dof	χ^2 -prob. (p)	σ_p^{SI}	σ_p^{SD}
CMSSM	32.8/18	11%	2σ	—
NUHM1	31.1/23	12%	1σ	—
NUHM2	30.3/22	11%	1σ	—
SU(5)	32.4/23	9%	1σ	—
mAMSB	36.5/27	11%	2σ	—
sub-GUT	28.9/24	23%	3σ	1σ below
pMSSM11	22.1/20	33%	3σ	1σ part. below

DM constraints:

⇒ **micrOMEGAs** for relic density and DD cross sections

[2019]

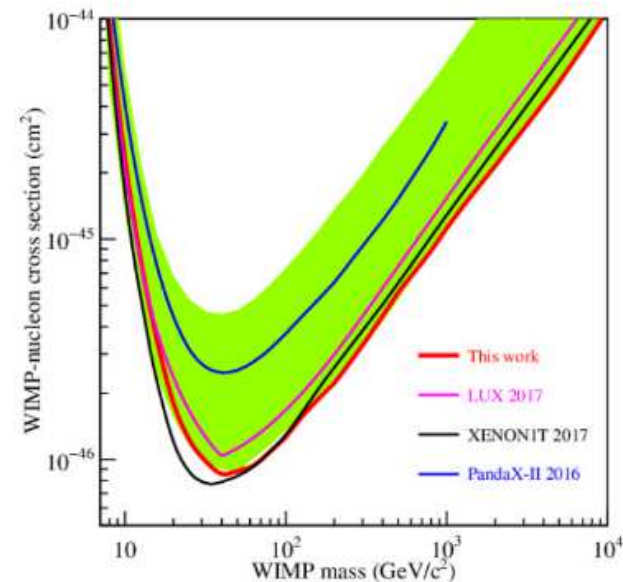
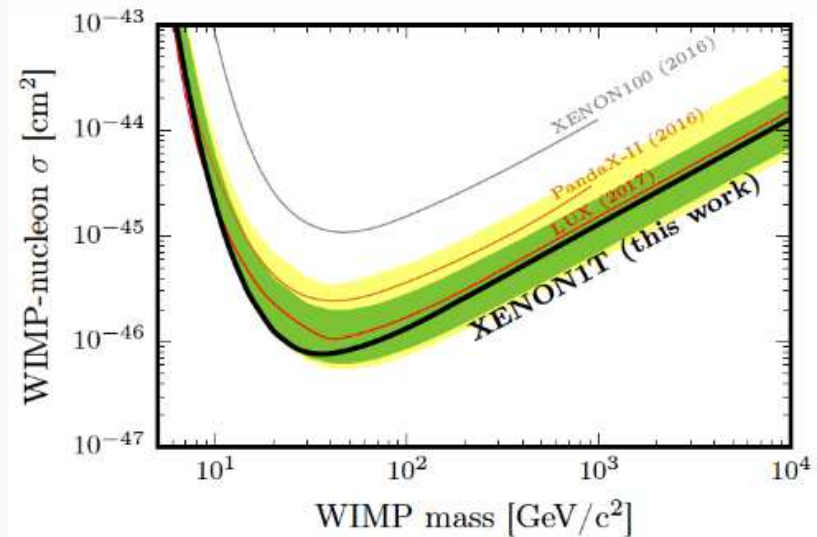
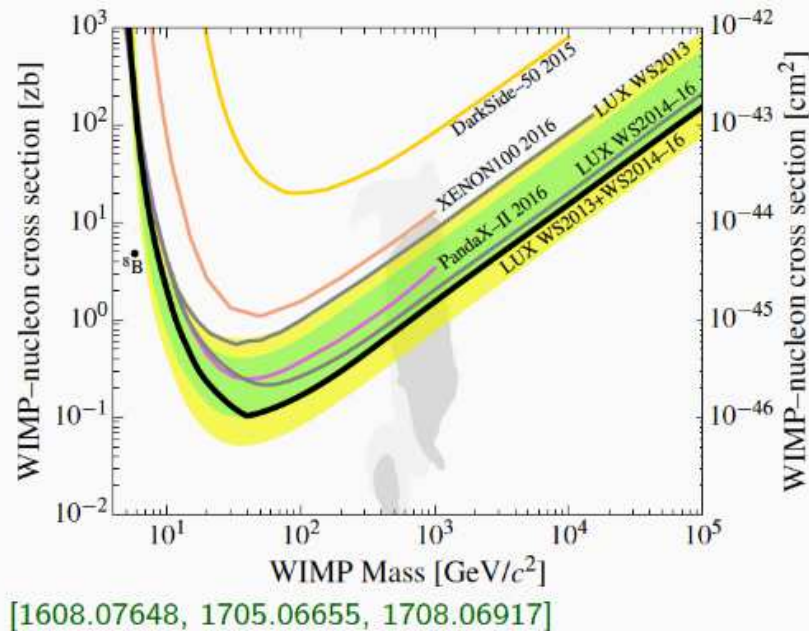


⇒ **full agreement** with ATLAS/CMS results (here: vector model)

Non-LHC constraints

Dark matter

- Relic density constraints from Planck.
- Direct detection constraints on σ_p^{SI} from LUX, XENON1T and PANDAX.
- Direct detection constraints on σ_p^{SD} from PICOD60.

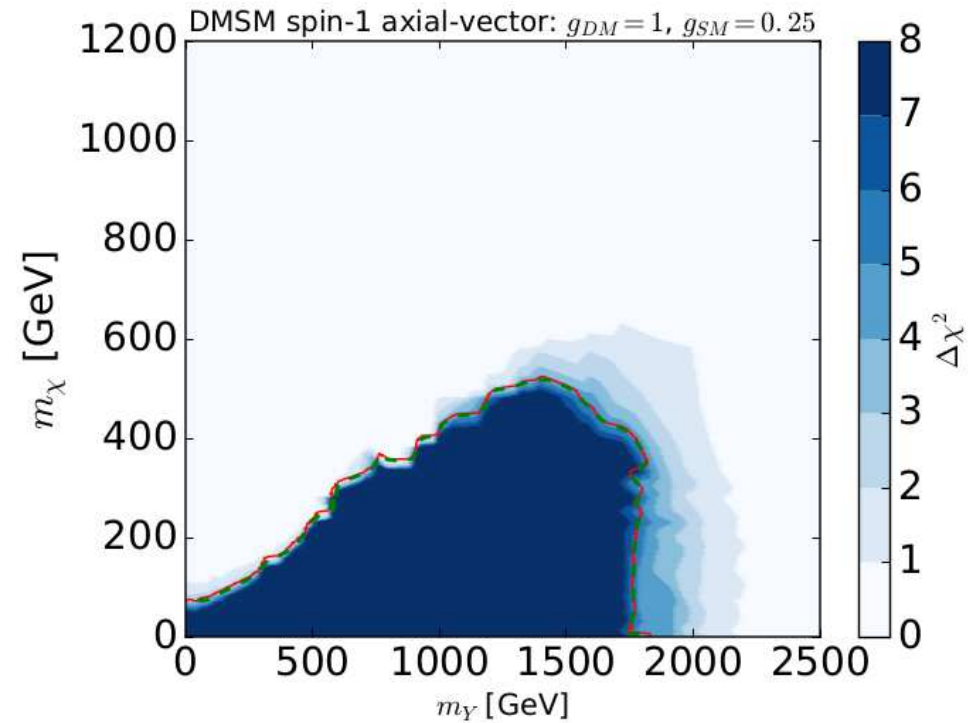
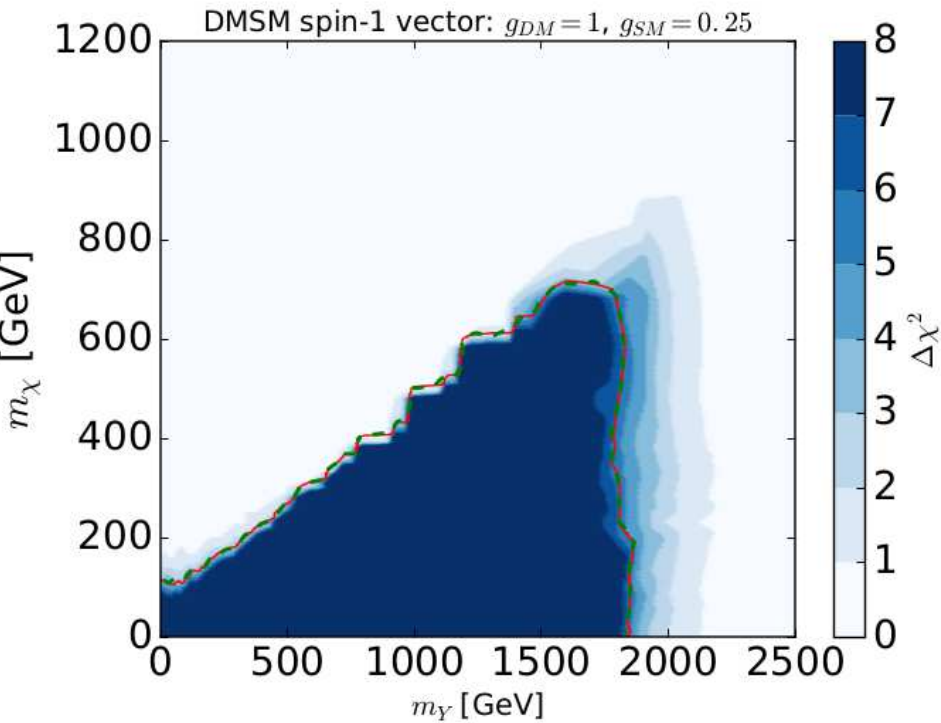


[taken from E. Bagnaschi]

Mono-jet constraints

[2019]

⇒ MG5 aMC(N)LO, Fastlim approach

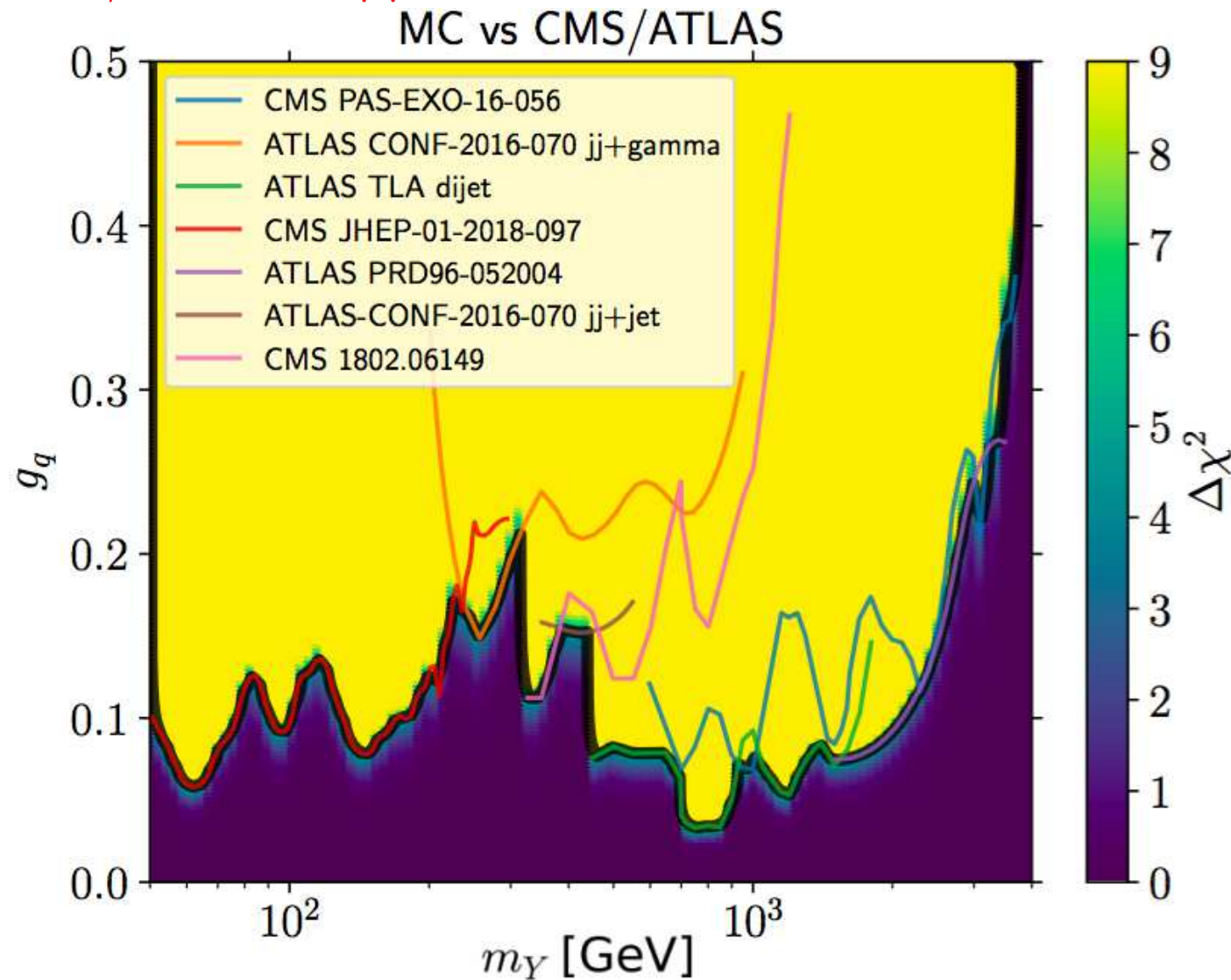


⇒ full agreement with ATLAS/CMS (red-dashed)

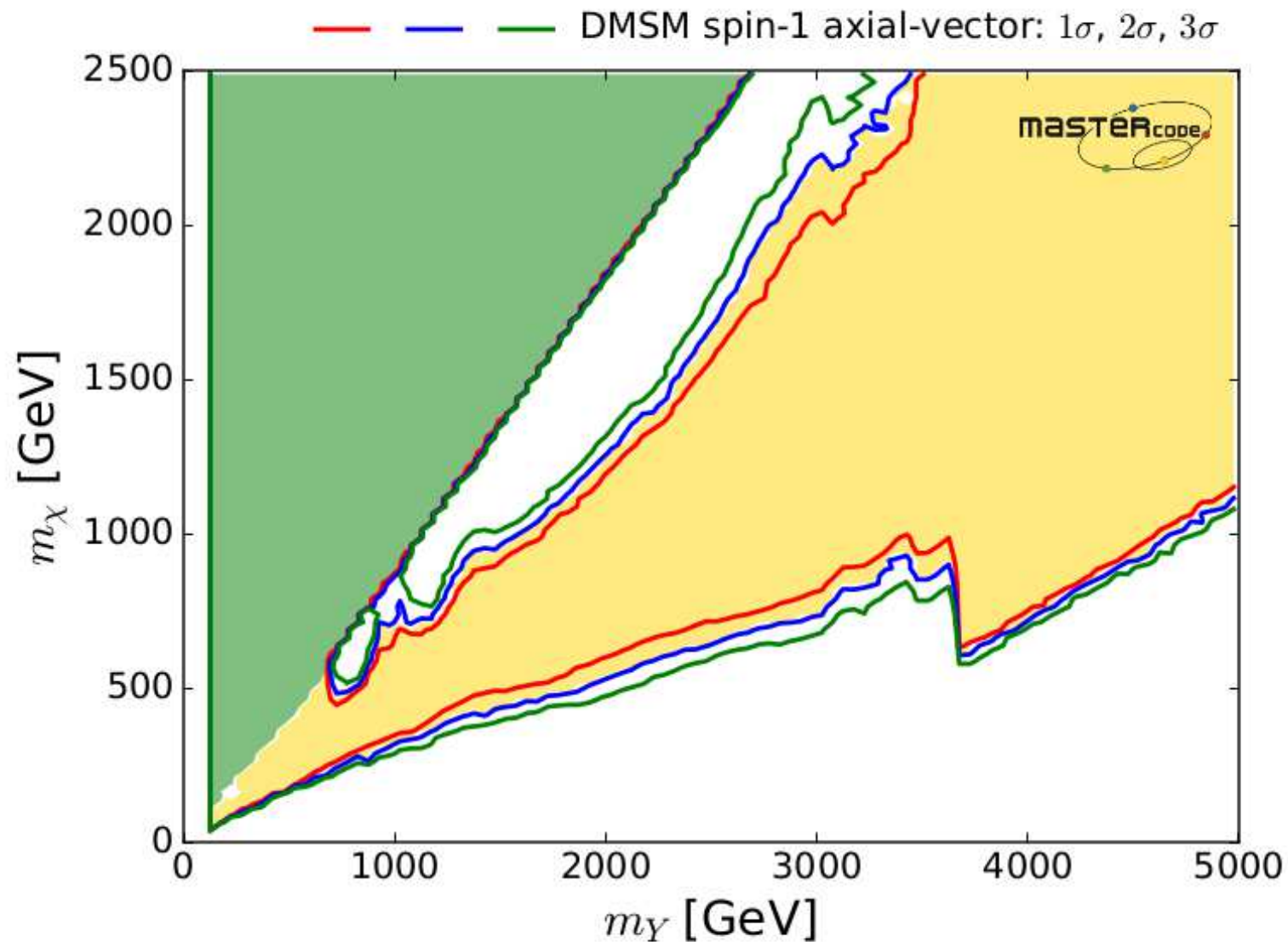
Di-jet constraints

⇒ MG5 aMC(N)LO, Fastlim approach

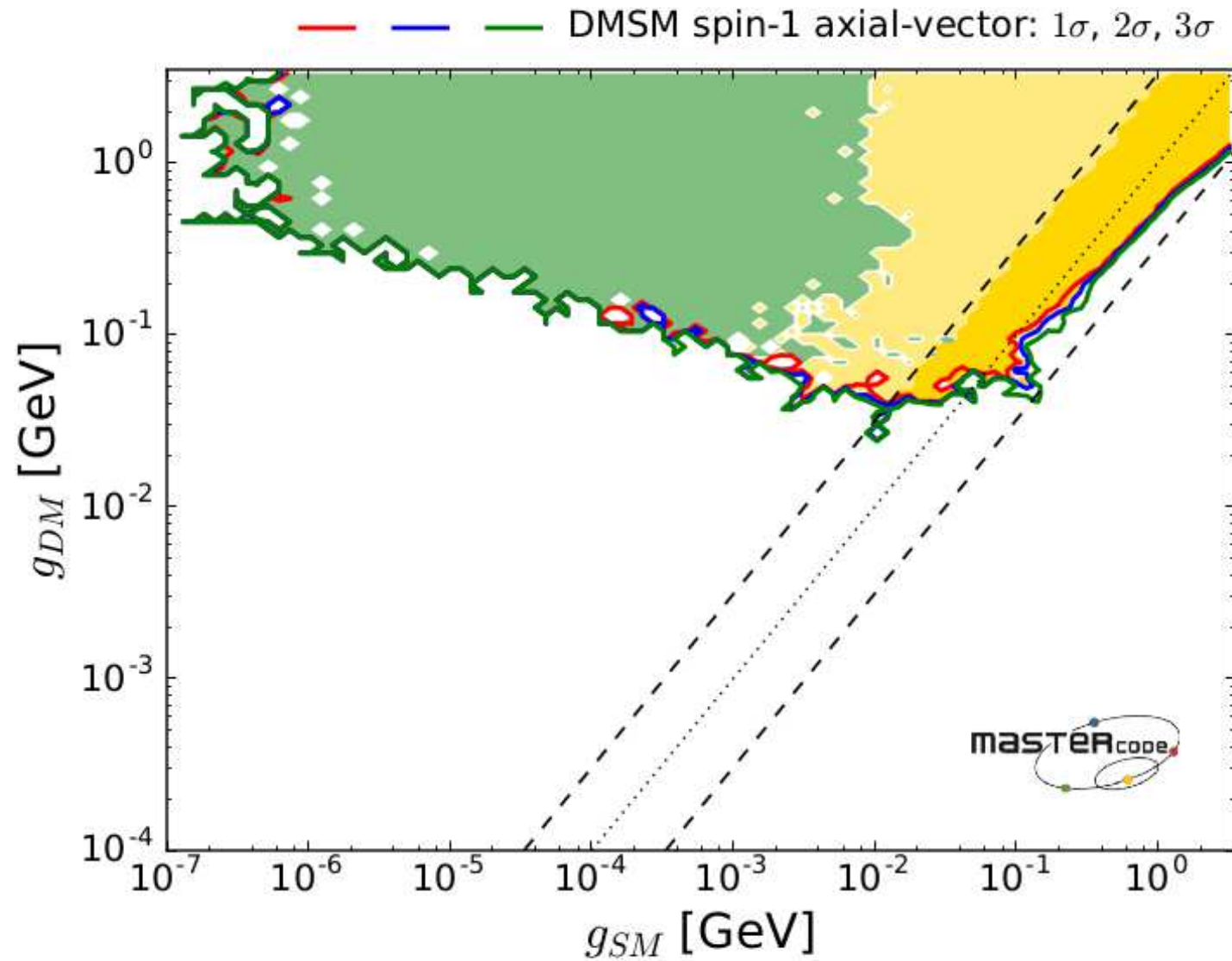
[2019]



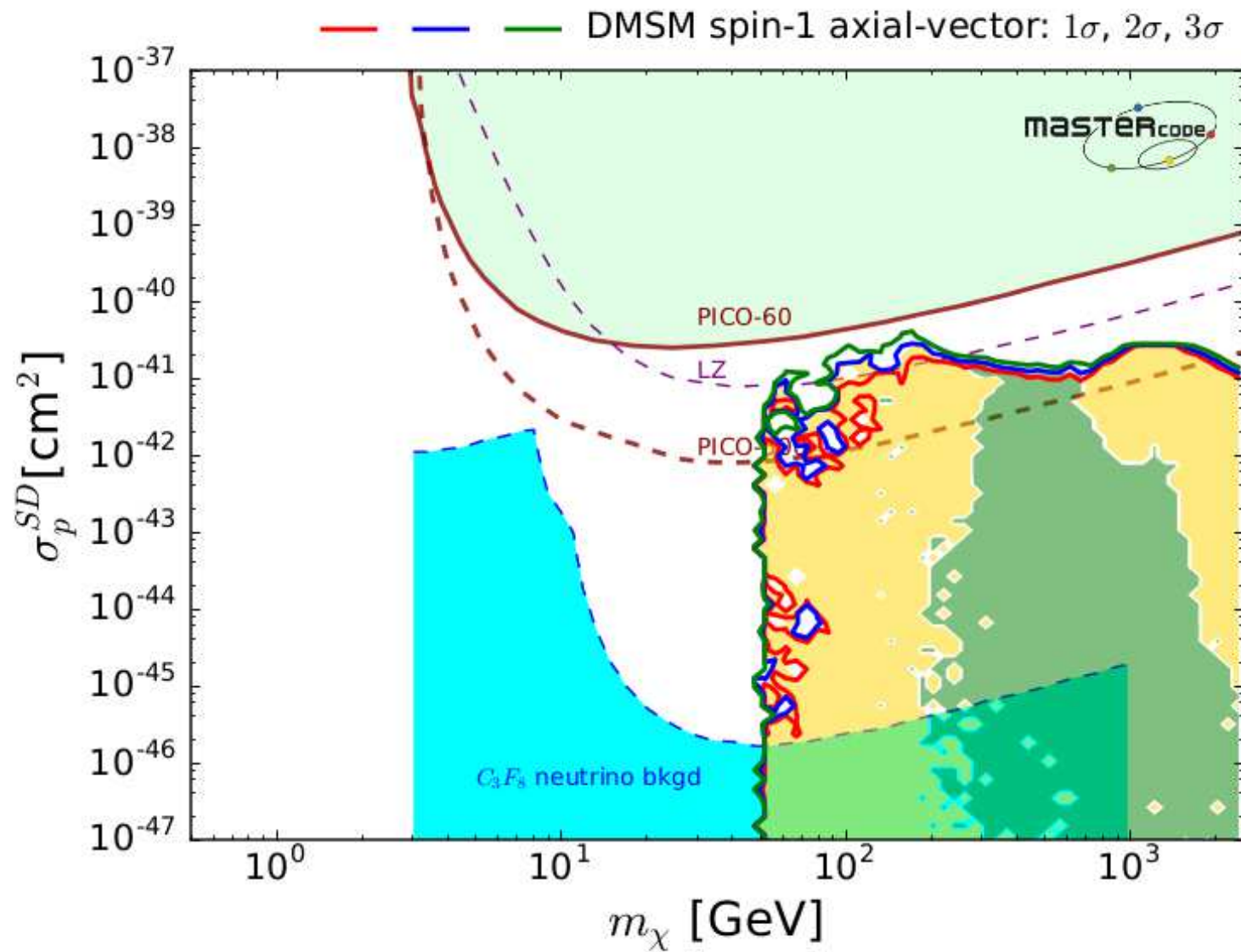
⇒ full agreement with ATLAS/CMS



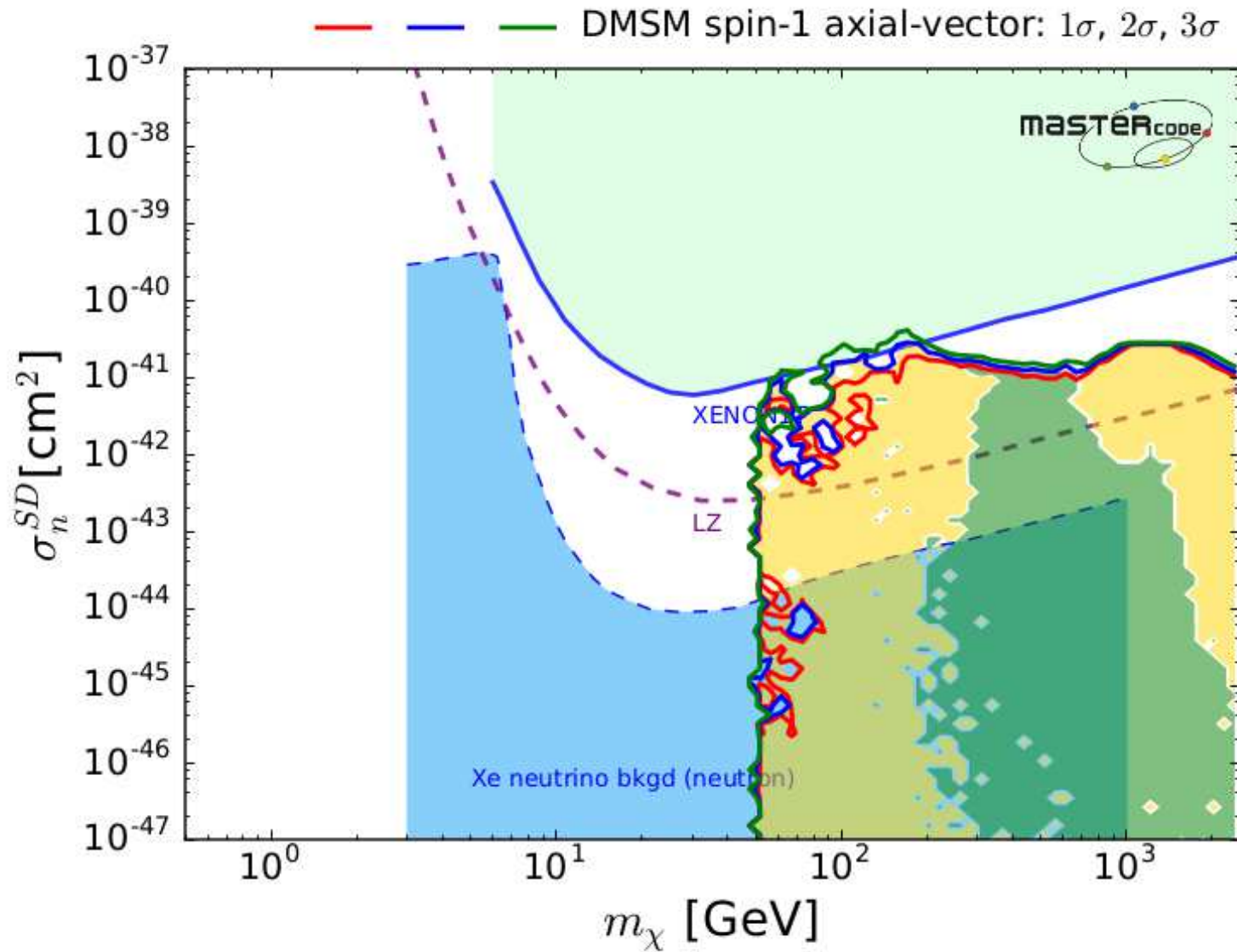
⇒ Larger s -channel region, continuous with t -channel



$\Rightarrow t$ - (s -)channel for $g_{SM} \lesssim (\gtrsim) 10^{-2}$



⇒ will not be easy for PICO!



⇒ neither for LZ!